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**Highlights**

- Global analysis of many rovibrational bands of CF<sub>4</sub>.
- This allows to simulate hot bands in the strongly absorbing 3 region.
- The results allow to model 92 % of the atmospheric absorption of this greenhouse gas.
- They have been used for spectroscopic database updates, including HITRAN2016.

# High-resolution spectroscopy and global analysis of CF<sub>4</sub> rovibrational bands to model its atmospheric absorption

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## Abstract

CF<sub>4</sub>, or tetrafluoromethane, is a chemically inert and strongly absorbing greenhouse gas, mainly of anthropogenic origin. In order to monitor and reduce its atmospheric emissions and concentration, it is thus necessary to obtain an accurate model of its infrared absorption. Such models allow opacity calculations for radiative transfer atmospheric models. In the present work, we perform a

global analysis (divided into two distinct fitting schemes) of 17 rovibrational bands of  $\text{CF}_4$ . This gives a reliable model of many of its lower rovibrational levels and allows the calculation of the infrared absorption in the strongly absorbing  $\nu_3$  region ( $1283 \text{ cm}^{-1} / 7.8 \text{ }\mu\text{m}$ ), including the main hot band, namely  $\nu_3 + \nu_2 - \nu_2$  as well as  $\nu_3 + \nu_1 - \nu_1$ ; we could also extrapolate the  $\nu_3 + \nu_4 - \nu_4$  absorption. This represents almost 92 % of the absorption at room temperature in this spectral region. A new accurate value of the C–F bond length is evaluated to  $r_e = 1.314860(21) \text{ \AA}$ . The present results have been used to update the HITRAN, GEISA and TFMecaSDa (VAMDC) databases.

*Keywords:* Carbon tetrafluoride, Tensorial Formalism, Spherical Top, Fourier Transform Infrared Spectroscopy, Global fit, Line Positions and Intensities

## 1. Introduction

Tetrafluoromethane ( $\text{CF}_4$ ), also known as Halocarbon 14, PFC-14 or R-14, is a colorless, non-toxic, non-flammable, non-corrosive gas belonging to the group of perfluorocarbons (PFCs). PFCs are extremely powerful greenhouse gases and find themselves among the longest-lived atmospheric traces gases because of their great chemical stability.  $\text{CF}_4$  is the most abundant PFC in the stratosphere, where it has an estimated lifetime of more than 50,000 years [1, 2]. For these reasons,  $\text{CF}_4$  is cited in the Kyoto Protocol [3, 4, 5] as a substance whose atmospheric concentration should be monitored and reduced. Infrared high spectral resolution solar occultation spectrometer measurements (Atmospheric Chemistry Experiment) led to an abundance of  $70.45 \pm 3.40$  pptv ( $10^{-12}$  per unit volume) [6]. Part of this is of natural origin [7, 8, 9]. The main source of  $\text{CF}_4$ , however, is anthropogenic and related to aluminum refining and to semiconductor manufacturing [10]. Recent studies from the ground and from space [6, 11, 12] detailed the evolution of  $\text{CF}_4$ 's atmospheric concentration.

The *in situ* detection of atmospheric  $\text{CF}_4$  is based on infrared solar absorption spectroscopy. The very intense  $\nu_3$  vibration-rotation band around  $7.8 \text{ }\mu\text{m}$  was used for the first detection of  $\text{CF}_4$  in the atmosphere from balloon-borne

observations [13]. Volumetric mixing ratio profiles of  $\text{CF}_4$  were retrieved from  
 20 satellite-borne [6, 14, 15] and balloon-borne observations [16]. The public spectroscopic databases like HITRAN [17, 18] or GEISA [19, 20], however, are not complete enough for atmospheric applications, since they completely lack hot band lines. This causes huge errors in  $\text{CF}_4$  concentration retrievals when using these databases, as shows in Ref. [21]. The authors of this paper already pointed  
 25 out that the lack of hot band lines in the databases is the main problem; thus, they used an empirical list of pseudo lines instead, which is much less flexible for simulations at any temperature.

Table 1 recalls the normal modes of this molecule. Using a simple harmonic approximation, it is easy to evaluate the population  $P(v_1, v_2, v_3, v_4)$  of the different vibrational levels, as a function of temperature  $T$  [22] as:  
 30

$$P(v_1, v_2, v_3, v_4) = \frac{g_v}{Q_v(T)} e^{-\frac{1}{k_B T}(v_1\tilde{\nu}_1 + v_2\tilde{\nu}_2 + v_3\tilde{\nu}_3 + v_4\tilde{\nu}_4)}, \quad (1)$$

where  $g_v$  is the vibrational level's degeneracy and  $Q_v(T)$  the vibrational partition function. This is plotted for  $\text{CF}_4$  on Figure 1. We thus see that the sole  $\nu_3$  fundamental band, also very strong, represents less than two-thirds of the absorption intensity at 296 K. In order to obtain a correct modelling of the molecule's opacity, it is thus mandatory to model hot bands, at least the one  
 35 starting from the first excited vibrational level,  $v_2 = 1$  ( $435.38 \text{ cm}^{-1}$ ), that is the  $\nu_3 + \nu_2 - \nu_2$  band.

Mode $\nu_i$	$\nu_1$	$\nu_2$	$\nu_3$	$\nu_4$
Symmetry	$A_1$	$E$	$F_2$	$F_2$
Degeneracy	1	2	3	3
Type	Stretching	Bending	Stretching	Bending
Activity	Raman	Raman	Absorption/Raman	Absorption/Raman
$\tilde{\nu}_i / \text{cm}^{-1}$	909.1	435.4	1283.2	631.1

Table 1: The normal modes of vibration of  $\text{CF}_4$ . Symmetry is the mode's irreducible representation in the  $T_d$  point group. Activity is given for the fundamental band.  $\tilde{\nu}_i$  is the approximate wavenumber for  $^{12}\text{CF}_4$  taken from the present study (from Tables A.6 and B.7).

To achieve this goal, it is thus necessary to know at first the lowest vibrational states (like bands with  $v_2 = v_3 = 1$  in the above example). The purpose of

40 the present paper is to perform a global and consistent analysis of as many  
rovibrational transitions as possible, gathering existing and new experimental  
data, in order to obtain a reliable modeling of  $\text{CF}_4$ 's absorption spectrum in the  
 $\nu_3$  region, including the most significant hot bands. We complete here our two  
previous studies [23] and [24] with extensive new experimental data, strongly  
45 increasing the number of assigned lines and of studied bands.

Section 2 details the various data used, while Section 3 recalls the spherical-  
top model used for analysis and modeling. In Section 4, we present line assign-  
ments and effective Hamiltonian parameter fits. Line intensities are discussed in  
Section 5 and, finally, Section 6 presents the recent database updates concerning  
50  $\text{CF}_4$ .

## 2. Experimental Details

This work gathers different data sources which are summarized in Table 2.  
We describe here the new spectra that have been recorded recently for this  
55 study.

It should be noticed that the uncertainty on line positions associated to high  
resolution Fourier Transform spectra is typically better than  $0.001 \text{ cm}^{-1}$ , while  
that of Raman spectra may be somewhat worse. However, knowing that we  
use here a huge number of lines from various sources, with different pressures,  
60 signal-to-noise ratios and apodization functions, we considered in the present  
study that all the line positions have a similar accuracy.

### 2.1. Infrared absorption data from Reims

The spectra are recorded with the Fourier Transform Spectrometer (FTS) of  
the GSMA laboratory already described elsewhere [25, 26]. The different char-  
65 acteristics of the spectra are reported in Table 2 (experimental conditions). The  
value of the diaphragm is 4 mm. The temperature of the cell was not stabilized,  
this is a 2 m-long White cell with an equivalent path length of 8.262 m. We

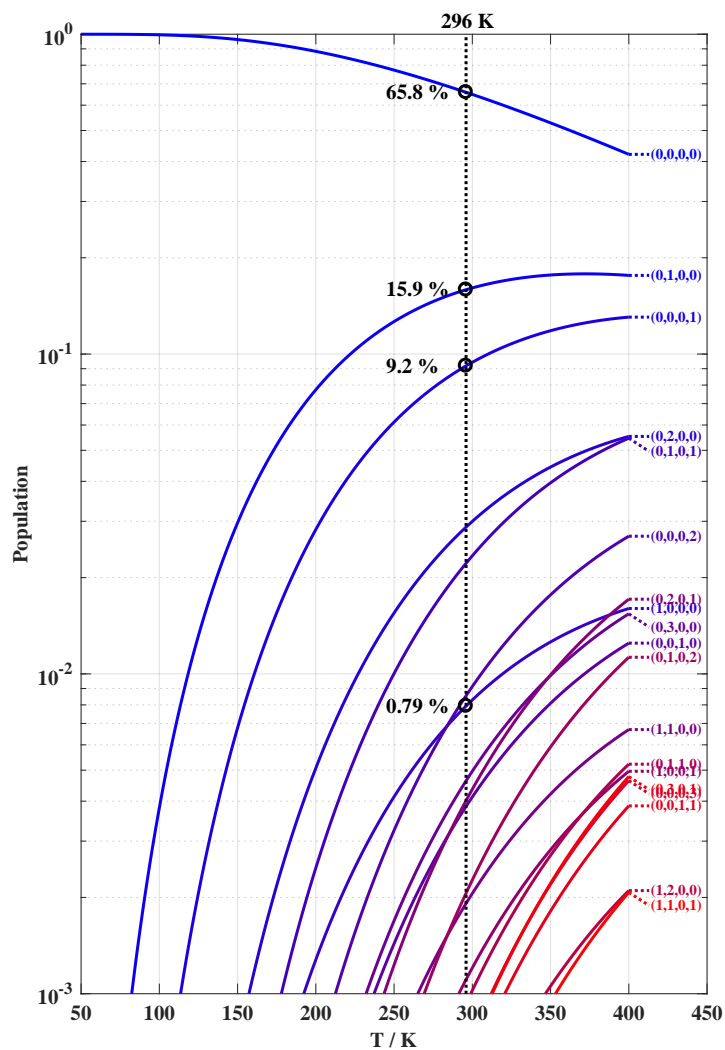


Figure 1: Population of the vibrational levels of CF<sub>4</sub>, as a function of temperature. Values at 296 K for the four lower levels involved in the hot bands calculated in the present work (see Section 5) are indicated. On the right, vibrational levels are labelled using the vibrational quantum number quadruplet,  $(v_1, v_2, v_3, v_4)$ , see Table 1 and Section 3.

use two different detectors, a photovoltaic HgCdTe (or MCT) detector cooled at 77 K for the  $\nu_2 + (\nu_3/2\nu_4)$  region and an InSb detector cooled at 77 K for the  $\nu_1 + (\nu_3/2\nu_4)$  region, a Globar source, and different selected optical filters.

### 2.2. Infrared absorption data from Braunschweig

Room temperature  $\text{CF}_4$  spectra were recorded with a high resolution FTS (Bruker IFS120HR) using a White multiple pass absorption cell with a path length of 4 m, equipped with KBr windows. The  $\nu_2 + \nu_4$  and  $\nu_1 + \nu_4$  bands were recorded using slightly different experimental conditions listed hereafter and grouped in Table 2.

A liquid nitrogen cooled MCT detector and a globar source combination was systematically used for the recording of the two bands. A KBr beam splitter was used for both bands. Gas pressures were monitored with a MKS capacitance gauge. The spectral resolution was  $0.003 \text{ cm}^{-1}$ . The  $\nu_1 + \nu_4$  band was monitored at a gas pressure of 15 Torr (20 mbar) by co-adding 70 scans while the  $\nu_2 + \nu_4$  band was recorded at a gas pressure of 11.25 Torr (15 mbar) by co-adding 200 scans.

### 2.3. Far infrared room temperature absorption data from SOLEIL

The far infrared room temperature data were recorded using the synchrotron radiation delivered by the AILES beamline of the synchrotron SOLEIL. The high-resolution FTS (Bruker IFS 125 HR) available on the beamline was connected to a White multiple pass optical cell to achieve an absorption path length of 150 m. The FTS was settled to its maximal resolution ( $0.00102 \text{ cm}^{-1}$ ) and a  $6 \mu\text{m}$  thick mylar beam splitter and a bolometer equipped with an internal optical filter were used to record the  $40\text{--}580 \text{ cm}^{-1}$  spectral range. A total of 326 and 274 scans were co-added for a gas pressure of 2.6 and 18 Torr respectively.

### 2.4. Jet-cooled infrared absorption data from SOLEIL

The jet-cooled absorption spectrum was recorded using the Jet-AILES apparatus implemented on the AILES beamline [27]. The infrared light beam



produced by the internal Globar source of the FTS was focused through a continuous planar supersonic expansion formed by a  $30\ \mu\text{m} \times 60\ \text{mm}$  slit nozzle. A throughput of about 6 slm (standard liter per minute) of  $\text{CF}_4$  in 20 slm of argon was regulated by appropriate Bronkhorst mass flow controllers all along the recording. The total stagnation pressure was 1530 Torr and the vacuum chamber pressure was 0.66 Torr. KBr windows were used to isolate the vacuum chamber from the FTS and the detector compartment see Ref. [27]. A combination of a KBr beam splitter, a MCT photovoltaic detector and an adapted optical filter was used to co-add 386 scans in the  $900\text{--}2700\ \text{cm}^{-1}$  spectral range, at a resolution of  $0.002\ \text{cm}^{-1}$ , and a rotational temperature of about 50 K.

### 2.5. THz rotational absorption data from SOLEIL

We also used newly recorded rotational lines in the  $v_3 = 1$  state ( $\nu_3 - \nu_3$  hot band) in the THz region ( $20\text{--}36\ \text{cm}^{-1}$ ). This dataset will be described in a forthcoming paper [28].

### 2.6. Earlier infrared absorption and Raman scattering data

We also use  $\nu_3/2\nu_4$  infrared absorption and  $\nu_1$ ,  $2\nu_2 - \nu_1$  and  $\nu_2$  stimulated Raman data taken from our previous works and described in Refs. [23] and [24], respectively.

## 3. Theoretical model

$\text{CF}_4$ , just like  $\text{CH}_4$  and other tetrahedral spherical top molecules with  $T_d$  point group symmetry at equilibrium, possesses four normal modes of vibration: one non-degenerate mode with  $A_1$  symmetry ( $\nu_1$ ), one doubly-degenerate mode with  $E$  symmetry ( $\nu_2$ ), and two triply-degenerate modes with  $F_2$  symmetry ( $\nu_3$  and  $\nu_4$ ), see Table 1. Only  $F_2$  fundamentals are infrared active, at first approximation, but other vibrational levels can gain some absorption intensity through Coriolis and anharmonic couplings (see for instance [29, 30]).

Transitions	Type	$\tilde{\nu}_0/\text{cm}^{-1}$	Place	$T/\text{K}$	$P/\text{Torr}$	$d/\text{m}$	$R/\text{cm}^{-1}$	Reference
Spectra used for the “ $\nu_2$ ” fit								
$\nu_3 - \nu_3$	THz	35	SOLEIL	296	100	150	0.001	[28]
$\nu_2$	IR	435	SOLEIL	296	18	150	0.001	This work
$\nu_2$	IR	435	IPR Rennes	296	501	970	0.02	[29] <sup>†</sup>
$\nu_2$	Raman	435	CSIC Madrid	296	1.5	—	0.003	[24]
$\nu_4$	IR	435	SOLEIL	296	2.6	150	0.001	This work
$\nu_2 + \nu_4$	IR	1066	Braunschweig	296	11.25	4	0.003	This work
$2\nu_2$	Raman	870	CSIC Madrid	135	22.5	—	0.003	[24]
$2\nu_2$	Raman	870	CSIC Madrid	296	11.3	—	0.003	[24]
$\nu_3/2\nu_4$	IR	1283	Braunschweig	Several spectra, see Reference				[23]
$\nu_3/2\nu_4$	IR jet	1283	SOLEIL	50	—	0.06	0.002	This work
$\nu_2 + (\nu_3/2\nu_4)$	IR jet	1720	SOLEIL	50	—	0.06	0.002	This work
$\nu_2 + (\nu_3/2\nu_4)$	IR	1720	GSMA Reims	295	5	8.262	0.003	This work
Spectra used for the “ $\nu_1$ ” fit								
$\nu_1 - \nu_4$	Far-IR	278	SOLEIL	296	18	150	0.001	This work
$\nu_1$	Raman	909	CSIC Madrid	135	1.5	—	0.003	[24]
$\nu_1$	Raman	909	CSIC Madrid	295	3.75	—	0.003	[24]
$2\nu_1 - \nu_1$	Raman	906	CSIC Madrid	296	22.5	—	0.003	[24]
$\nu_1 + \nu_4$	IR	1540	Braunschweig	296	15	4	0.003	This work
$\nu_1 + (\nu_3/2\nu_4)$	IR jet	2186	SOLEIL	50	—	0.06	0.001	This work
$\nu_1 + (\nu_3/2\nu_4)$	IR	2186	GSMA Reims	295	0.5	8.262	0.003	This work

<sup>†</sup> This spectrum was not used for the present analyses, but it is compared to a high-pressure simulation in Figure 5 below.

Table 2: Experimental spectra used for the analyses. They were recorded in different laboratories.  $\tilde{\nu}_0$  is the approximate wavenumber,  $T$  the temperature,  $P$  the pressure,  $d$  the optical path and  $R$  the spectral resolution.

### 3.1. Effective Hamiltonian

In this paper we use the theoretical model based on the tensorial formalism and the vibrational extrapolation concept developed by the Dijon group [31, 32]. It takes full advantage of the molecule’s high symmetry. Let us just recall briefly the principles of this model which have been already detailed in Ref. [32]. Considering an  $\text{XY}_4$  molecule such as  $\text{CF}_4$ , the vibrational levels are grouped in series of polyads named  $P_k$  with  $k = 0, \dots, n$ . For  $k = 0$ , we have  $P_0$  which is the ground state ( $GS$ ). The Hamiltonian operator is written as follows:

$$\mathcal{H} = \mathcal{H}_{\{P_0 \equiv GS\}} + \mathcal{H}_{\{P_1\}} + \dots + \mathcal{H}_{\{P_k\}} + \dots + \mathcal{H}_{\{P_{n-1}\}} + \mathcal{H}_{\{P_n\}}. \quad (2)$$

where the different  $\mathcal{H}_{\{P_k\}}$  terms are expressed in the following form:

$$\mathcal{H}_{\{P_k\}} = \sum_{\text{all indexes}} t_{\{s\}\{s'\}}^{\Omega(K,n\Gamma)\Gamma_v\Gamma'_v} \beta \left[ \varepsilon V_{\{s\}\{s'\}}^{\Omega_v(\Gamma_v\Gamma'_v)\Gamma} \otimes R^{\Omega(K,n\Gamma)} \right] (A_1). \quad (3)$$

In this equation, the  $t_{\{s\}\{s'\}}^{\Omega(K,n\Gamma)\Gamma_v\Gamma'_v}$  are the parameters to be determined, while  $\varepsilon V_{\{s\}\{s'\}}^{\Omega_v(\Gamma_v\Gamma'_v)\Gamma}$  and  $R^{\Omega(K,n\Gamma)}$  are vibrational and rotational operators, respectively.

For each term,  $\Omega_v$  and  $\Omega$  represent the degree in elementary vibrational operators (creation  $a^+$  and annihilation  $a$  operators), and rotational operators (components  $J_x, J_y$  and  $J_z$  of the angular momentum), respectively.  $\beta$  is a factor that allows the scalar terms (terms with  $\Gamma = A_1$ , the totally symmetric irreducible representation of  $T_d$ ) to match the “usual” contributions like  $B_0 J^2$ , etc. The order of each individual term is defined as  $\Omega + \Omega_v - 2$ . We deal with the effective Hamiltonians which are obtained, for a given polyad  $P_k$ , by the projection of  $H$  in the  $P_n$  Hilbert subspace:

$$\begin{aligned} H^{<P_n>} &= P^{<P_n>} \mathcal{H} P^{<P_n>} \\ &= H_{\{\text{GS}\}}^{<P_n>} + H_{\{P_1\}}^{<P_n>} + \dots + H_{\{P_k\}}^{<P_n>} + \dots + H_{\{P_{n-1}\}}^{<P_n>} + H_{\{P_n\}}^{<P_n>}. \end{aligned} \quad (4)$$

This effective Hamiltonian contains all the operators, and thus all the parameters, of the lower polyads. Parameters of upper polyads are (presumably) small corrections to those of the lower ones since in our model they are in fact differences between lower and upper state values; this is what we call “vibrational extrapolation”. It enables global fits of data for a series of consecutive polyads, ensures a global consistency and, usually, a better convergence of molecular parameters. Just in order to give a simple illustrative example, if we consider a polyad  $P_1$  containing a fundamental band, then the effective Hamiltonian  $H^{<P_1>} = H_{\{\text{GS}\}}^{<P_1>} + H_{\{P_1\}}^{<P_1>}$  contains in  $H_{\{\text{GS}\}}^{<P_1>}$  the ground state rotational constant  $B_0$  and in  $H_{\{P_1\}}^{<P_1>}$  the difference  $\Delta B = B_1 - B_0$ , that is the difference between the upper (say,  $B_1$ ) and lower ( $B_0$ ) state rotational constants, and not  $B_1$  itself. We usually have  $\Delta B \ll B_0, B_1$ .

### 3.2. Polyad scheme definition

Generally speaking, the polyad number  $n$  can be defined as

$$n = i_1 v_1 + i_2 v_2 + i_3 v_3 + i_4 v_4, \quad (5)$$

where  $v_1, v_2, v_3$  and  $v_4$  are vibrational quantum numbers for the four normal modes of the molecule, while  $i_1, i_2, i_3$  and  $i_4$  are integer numbers chosen to group vibrational levels  $(v_1, v_2, v_3, v_4)$  into polyads, according to their wavenumbers.

In particular, when some vibrational numbers display simple approximate ratios, a clear polyad structure may appear. For instance, the well-known polyad structure of methane [33] comes from the fact that  $\tilde{\nu}_1 \simeq \tilde{\nu}_3 \simeq 2\tilde{\nu}_2 \simeq 2\tilde{\nu}_4$ , leading to large absorbing bands which are regularly spaced, each  $1500 \text{ cm}^{-1}$ . In the case of  $\text{CF}_4$ , however, no such global approximate relation exists. In Ref. [23], it was just noticed that  $\tilde{\nu}_3 \simeq 2\tilde{\nu}_4$ , implying to place the  $v_3 = 1$  and  $v_4 = 1$  vibrational levels in the same polyad. Thus, contrary to what is done for  $\text{CH}_4$ , it appears difficult to include all modes in a single polyad scheme at this stage. We thus decided here to build two separate polyad schemes:

- A “ $\nu_2$ ” polyad scheme implying the  $\nu_2$  bending mode, its overtones and combinations and related bands linked to it through combinations (in fact, almost “all” the lowest  $\text{CF}_4$  bands not linked with the  $\nu_1$  mode). In this case, we choose:

$$n_{\nu_2} = 2v_2 + 6v_3 + 3v_4. \quad (6)$$

- A “ $\nu_1$ ” polyad scheme implying the  $\nu_1$  stretching mode influence into several bands. In this case, we choose:

$$n_{\nu_1} = 3v_1 + 4v_3 + 2v_4. \quad (7)$$

In both cases, there is no polyad number 1 (this one has no corresponding vibrational level). As we will see later, these two schemes lead to two separate global effective Hamiltonian fits. Tables 3 and 4 list essential details on the polyads including the number of operators for the “ $\nu_2$ ” and “ $\nu_1$ ” schemes, respectively. All vibrational quantum numbers are limited to the values 0, 1 and 2. Except for the (non-existing) polyad number 1, all effective Hamiltonian contributions have been expanded up to the 8th order. In the final analysis, however (see Section 4), only a small subset of all the resulting parameters could be actually determined from the fit, as indicated in the last column of both tables.

Figure 2 shows the transitions included in the two fits, corresponding to the two polyad schemes. These fits are described in Section 4 below.

Polyad	Vibrational levels	Order	Nb. parameters	Nb. fitted
0	(0,0,0,0)	8	15	10
1	—	0	0	0
2	(0, 1, 0, 0)	8	26	12
3	(0, 0, 0, 1)	8	48	17
4	(0, 2, 0, 0)	8	29	6
5	(0, 1, 0, 1)	8	84	9
6	(0, 0, 1, 0), (0, 0, 0, 2)	8	224	51
7	(0, 2, 0, 1)	8	78	0
8	(0, 1, 1, 0), (0, 1, 0, 2)	8	383	92
Total			887	197

Table 3: Effective Hamiltonian for the “ $\nu_2$ ” scheme, with the number of parameters in the model. Vibrational levels are denoted using vibrational quantum numbers in the form  $(v_1, v_2, v_3, v_4)$ . The last column shows the number of parameters that could be fitted (see Section 4.1).

### 3.3. Line intensity calculations

185 In order to calculate transition intensities, the effective dipole moment transition operator is also determined in a similar way (see Ref. [32] for more detail). This operator is expanded to the lowest possible degree. In the case of the  $\nu_3/2\nu_4$  dyad, we use the same dipole moment expansion as in our previous paper [23]. Dipole moment derivative values and line intensities in this strongly absorbing  
190  $\nu_3$  region are discussed in Section 5.

### 3.4. Basis sets

The calculation of the effective Hamiltonian and effective dipole moment matrix elements are performed in the coupled rovibrational basis

$$\left| \left[ \Psi_v^{(C_v)} \otimes \Psi_r^{(J, nC_r)} \right]^{(C)} \right\rangle, \quad (8)$$

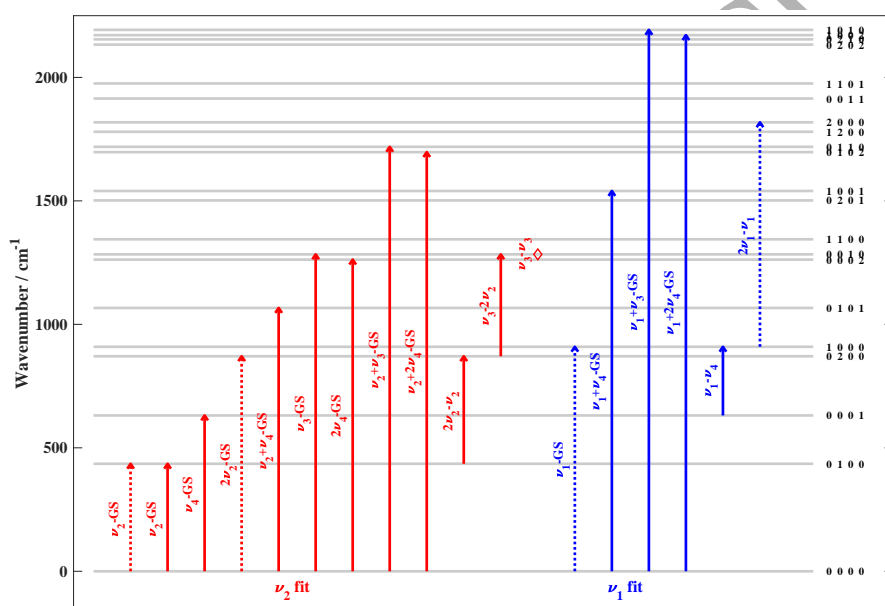


Figure 2: Vibrational transitions and levels used in the two fits. Dashed lines correspond to Raman transitions. The  $\nu_2$  fundamental band is observed both in infrared absorption and Raman scattering (see text).

Polyad	Vibrational levels	Order	Nb. parameters	Nb. fitted
0	(0,0,0,0)	8	15	0 <sup>†</sup>
1	—	0	0	0
2	(0, 0, 0, 1)	8	48	0 <sup>†</sup>
3	(1, 0, 0, 0)	8	11	7
4	(0, 0, 1, 0), (0, 0, 0, 2)	8	224	0 <sup>†</sup>
5	(1, 0, 0, 0)	8	27	24
6	(2, 0, 0, 0), (0, 0, 1, 1)	8	214	3 <sup>‡</sup>
7	(1, 0, 1, 0), (1, 0, 0, 2)	8	109	73
Total			648	107

<sup>†</sup> Parameters fixed to the values from the “ $\nu_2$ ” fit, see Table 3.

<sup>‡</sup> Only  $v_1 = 2$  parameters are fitted.

Table 4: Effective Hamiltonian for the “ $\nu_1$ ” scheme, with the number of parameters in the model. Vibrational levels are denoted using vibrational quantum numbers in the form  $(v_1, v_2, v_3, v_4)$ . The last column shows the number of parameters that could be fitted (see Section 4.2).

195 where  $\Psi_r^{(J, n C_r)}$  is a rotational wavefunction with angular momentum  $J$ , rotational symmetry species  $C_r$  and multiplicity index  $n$ ;  $\Psi_v^{(C_v)}$  is a coupled vibrational basis set;  $C$  is the overall symmetry species ( $C = C_v \otimes C_r$ ).  $\Psi_v^{(C_v)}$  contains the relevant functions for the 4 normal modes of vibration,

$$\left| \Psi_v^{(C_v)} \right\rangle = \left| \left( (\Psi_{v_1}^{(A_1)} \otimes \Psi_{v_2}^{(l_2, C_2)} \otimes \Psi_{v_3}^{(l_3, n_3 C_3)})^{(C_{23})} \otimes \Psi_{v_4}^{(l_4, n_4 C_4)} \right)_{\sigma_v}^{(C_v)} \right\rangle \quad (9)$$

For each normal mode number  $i = 1$  to 4, we use a symmetrized harmonic  
200 oscillator basis set denoted

$$\left| \psi_{v_i}^{(l_i, n_i C_i)} \right\rangle = |v_i, l_i, n_i C_i, \sigma_i\rangle, \quad (10)$$

where

$$l_1 = 0, \quad (11)$$

$$l_i = v_i, v_i - 2, v_i - 4, \dots, 0 \text{ or } 1 \quad (i = 2, 3 \text{ or } 4), \quad (12)$$

are the usual vibrational angular momentum quantum numbers for degenerate vibrations.  $C_i$  is a  $T_d$  irreducible representation (irrep) and  $n_i$  a multiplicity index used when  $C_i$  appears more than once for a given  $(v_i, l_i)$  set. We always  
 205 have  $n_1 = 0$ ,  $n_2 = 0$  and necessarily  $C_1 = A_1$ . In Equation (9), we have:

$$C_{23} = C_2 \otimes C_3 \quad \text{and} \quad C = C_{23} \otimes C_4, \quad (13)$$

$C_v$  being the overall vibrational symmetry. All  $C$  symbols here represent  $T_d$  irreps.

#### 4. Global fits of line positions

As already mentioned, the philosophy of this work is to gather many  $\text{CF}_4$   
 210 high-resolution spectroscopic data (see Table 2) to use them in global fits of line positions. Just as explained in Section 3.2, it appears difficult to include all  $\text{CF}_4$  vibrational levels in a single polyad scheme. As a matter of fact, contrary to methane ( $\text{CH}_4$ ) which has a well-defined polyad structure [33, 34, 35],  $\text{CF}_4$  has fundamental vibrational wavenumber with a more random distribution; we  
 215 can just notice, as in our previous work [23], that  $\tilde{\nu}_3$  is close to  $2\tilde{\nu}_4$  leading to a Fermi interaction between the corresponding bands that must be treated as a dyad. This interaction propagates to all combination bands built on  $\nu_3/2\nu_4$ , for instance when adding  $\nu_1$  or  $\nu_2$  quanta.

Thus, it appeared wise and more practical to perform assignments and effective  
 220 Hamiltonian fit following two partial polyad schemes. These two schemes are described in Section 3.2 and the two sections below describe the assignment and fit process in each case.

It should be noticed that during the fits the different sets of data were used  
 225 with the same weights due to compatible quality of the experimental spectra (see Section 2).

##### 4.1. Global fit based on the $\nu_2$ mode

The fundamental band of the  $\nu_2$  mode of tetrahedral  $\text{XY}_4$  molecules is, at first approximation, only active in Raman scattering. A few years ago, we could



analyze this fundamental band thanks to high-resolution stimulated Raman spectroscopy [24]. This excellent technique can provide essential data for such non-infrared active modes [36, 37, 38, 39, 24] but has also some drawbacks: its spectral resolution (around  $0.003 \text{ cm}^{-1}$ ) and sensitivity are significantly lower than what can be achieved using infrared absorption spectroscopy. In some cases, however,  $\nu_2$  can gain intensity from rovibrational coupling with another vibrational mode. This is the case for methane, for instance, whose  $\tilde{\nu}_2$  and  $\tilde{\nu}_4$  wavenumbers are quite close to each other [40], so that the two corresponding fundamentals form a bending dyad. Although these two modes are more separated for  $\text{CF}_4$ , a weak  $\nu_2$  fundamental band is observable in the  $400\text{--}470 \text{ cm}^{-1}$  region when a long optical path length is used [29]. In other words, the weakness of this interaction does not necessitate to consider a  $\nu_2 = \nu_4 = 1$  polyad, but its influence on the  $\nu_2$  induced dipole moment is sufficient to let this fundamental band appear in infrared absorption. In the present work, we used such a spectrum (see Section 2 and Table 2), along with the existing Raman data from Ref. [24].

We then included new data concerning the  $\nu_2 + \nu_4$  combination band. Analyzing this one implies to know the  $\nu_4 = 1$  level itself, so the previous data on the  $\nu_4$  fundamental [23] were necessary. As this one was already a part of a global study including the  $\nu_3/2\nu_4$  dyad (see above), we also included data corresponding to this region, coming both from Ref. [23] and from newly recorded spectra (Table 2). Some recently recorded THz lines of the  $\nu_3 - \nu_3$  hot band are also used [28]. Finally, a new  $\nu_2 + (\nu_3/2\nu_2)$  spectrum was also considered, leading to a global analysis and fit of all rovibrational bands that do not imply the  $\nu_1$  mode in the  $0$  to  $1750 \text{ cm}^{-1}$  region. We also noticed that the new spectra allowed to assign some hot band lines for  $2\nu_2 - \nu_2$  and  $\nu_3 - 2\nu_2$  in the  $\nu_2$  infrared spectrum, which completed the dataset. As explained in Section 3.2 and Table 3, this is referred as the “ $\nu_2$ ” scheme and fit.

We can notice that this study allowed to refine the GS parameters. It was also necessary to correct some  $\nu_2$ ,  $2\nu_2$  and  $\nu_3/2\nu_4$  assignments from our previous papers [23, 24].  $\nu_4$  and  $\nu_3/2\nu_4$  are now assigned up to slightly higher  $J$  values.

260 The analysis was performed thanks to the STDS [41] package of the XTDS  
 software [42] which allows to calculate and fit  $XY_4$  spherical-top spectra. All  
 assignments have been performed with help of the SPVIEW complementary  
 software [42]. A total of 17,626 lines could be assigned (see Table 5) to determine  
 197 effective Hamiltonian parameters (see Table 3). The overall root mean  
 265 square deviation is  $1.116 \times 10^{-3} \text{ cm}^{-1}$ . For each of the new spectra, initial  
 simulations were performed using the existing effective Hamiltonian parameters  
 from the previous studies which yielded preliminary assignments. Then the new  
 parameters were determined and the assignments refined. Some corrections were  
 also made to the assignments from the previous works. The process was iterated  
 270 till no more assignment appeared possible.

Table 5 summarizes the fit statistics and Figure 3 displays the fit residuals  
 for line positions. Table A.6 in the Appendix gives the effective Hamiltonian  
 parameter values. Concerning the  $\nu_2$  fundamental band itself, it should be  
 noticed that we now have 3.2 times more lines, for  $J$  up to 82, instead of 45,  
 275 compared to Ref. [24].

We give four examples of comparisons between experiment and simulation  
 in spectral regions in infrared absorption, which are analyzed and modeled here  
 for the first time. Figure 4 presents the interaction-induced  $\nu_2$  region for which  
 280 we could analyse and simulate two hot bands, namely  $2\nu_2 - \nu_2$  and  $\nu_3 - 2\nu_2$ .  
 Figure 5 represents the same region at a high pressure, the experimental data  
 being those used for Figure 1 of Ref. [29]. For these first two examples, the  
 vertical scale of the simulations has been adjusted manually to match with the  
 experimental spectra. No attempts to determine absolute intensities has been  
 285 performed in this case, contrary to what is discussed in Section 5 below for the  
 strongly absorbing  $\nu_3$  region. Figure 6 displays the  $\nu_2 + \nu_4$  region. Figure 7  
 shows some details concerning the quite complex  $\nu_2 + (\nu_3/2\nu_4)$  region. This  
 one has 8 vibrational sublevels. We illustrate this on Figure 8, which displays  
 the rovibrational sublevels, as a function of the rotational quantum number  $J$ .

Transitions	Number of data	$J_{\max}$	$\sigma$	$d_{\text{RMS}} / 10^{-3} \text{ cm}^{-1}$
“ $\nu_2$ ” fit				
$\nu_2$	4686	82	0.9984	1.027
$\nu_4$	2254	72	1.276	1.276
$2\nu_2$	823	42	0.7757	0.7757
$\nu_2 + \nu_4$	877	39	1.188	1.188
$\nu_3/2\nu_4$	2930	53	0.9847	1.161
$\nu_2 + (\nu_3/2\nu_4)$	4763	49	1.167	1.167
$2\nu_2 - \nu_2$	803	68	0.7758	0.7758
$\nu_3 - 2\nu_2$	222	37	1.189	1.189
$\nu_3 - \nu_3$	268	37	0.9173	33.64 MHz
Total	17626	82	1.076	1.116
“ $\nu_1$ ” fit				
$\nu_1$	2368	83	0.3498	1.049
$\nu_1 + \nu_4$	669	47	0.4689	0.8687
$\nu_1 - \nu_4$	2524	64	1.043	1.043
$\nu_1 + (\nu_3/2\nu_4)$	3094	74	1.608	1.825
$2\nu_1 - \nu_1$	229	73	3.068	0.6968
Total	8884	83	1.225	1.354

Table 5: Fit statistics for the two polyad schemes.  $\sigma$  is the dimensionless standard deviation and  $d_{\text{RMS}}$  the root mean square deviation in  $10^{-3} \text{ cm}^{-1}$ , except for THz  $\nu_3 - \nu_3$  lines for which it is given in MHz.

290 More precisely, this Figure shows the reduced energy levels, defined by:

$$\begin{aligned}
\tilde{\nu}_{\text{red}} &= \tilde{\nu} - \sum_{\Omega} t_{\{\text{GS}\}\{\text{GS}\}}^{\Omega(0,0A_1)A_1A_1} (J(J+1))^{\Omega/2} \\
&= \tilde{\nu} - B_0 J(J+1) + D_0 J^2(J+1)^2 - \dots, \quad (14)
\end{aligned}$$

*i.e.* we subtract the dominant scalar polynomial terms in order to enhance levels splittings due to molecular symmetry. We give both the calculated and observed reduced energy levels. Observed levels are simply levels reached by assigned transitions which are included in the fit. This gives a good idea of the

295 sampling of the energy spectrum.

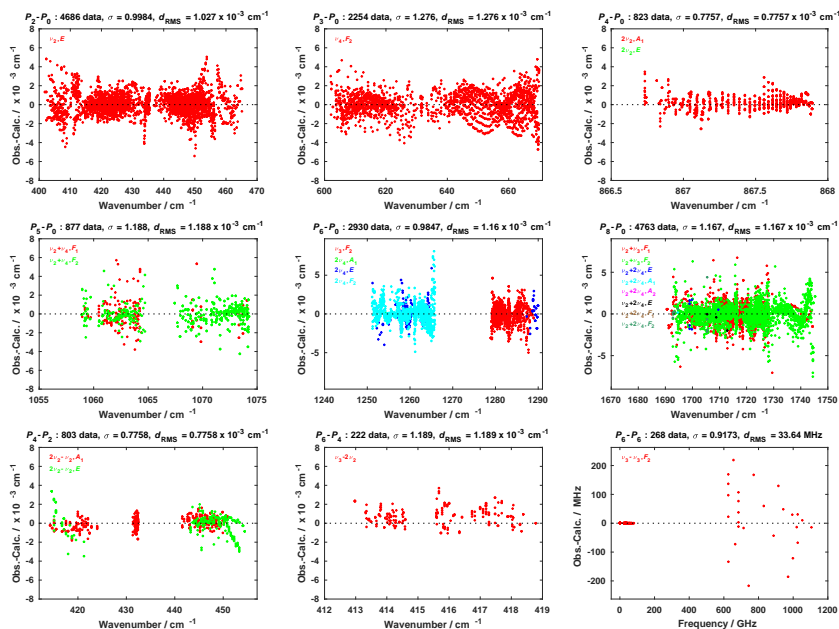


Figure 3: Fit residuals for line positions, in the case of the “ $\nu_2$ ” fit. In each spectral region, the colors correspond to the different vibrational sublevels.

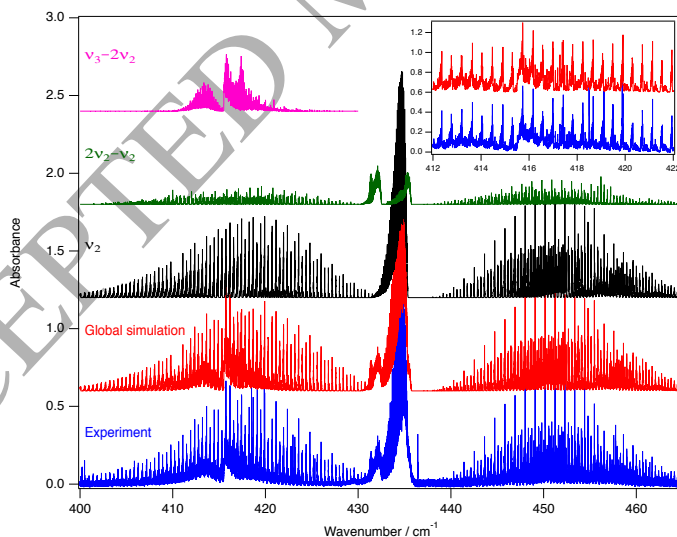


Figure 4: Measured  $\nu_2$  spectral region in infrared absorption at 296 K and 18 Torr, compared to the simulation with two hot bands. The insert details a part of the  $P$  branch, around the  $\nu_3 - 2\nu_2$  region.

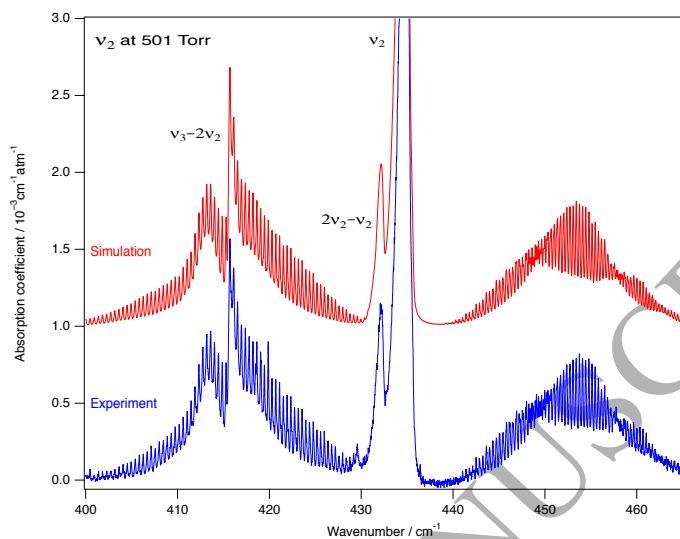


Figure 5: Measured  $\nu_2$  spectral region at 296 K and high pressure (501 Torr, see Table 2 and Ref. [29]), compared to the simulation with two hot bands.

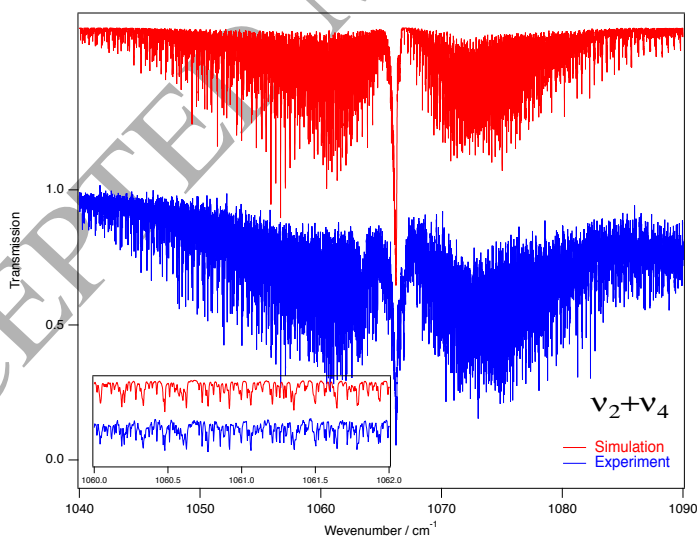


Figure 6: Measured  $\nu_2 + \nu_4$  difference band at 296 K and 11.25 Torr, compared to the simulation. The insert details a part of the  $P$  branch.

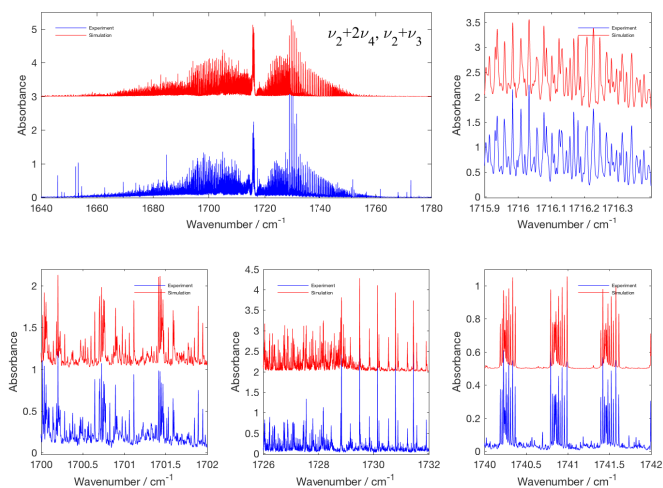


Figure 7: Measured spectrum in the  $\nu_2 + (\nu_3/2\nu_4)$  region (295 K, 5 Torr), compared to the simulation, with detailed view of some parts. Strong lines on the wings are water vapor lines.

300

#### 4.2. Global fit based on the $\nu_1$ mode

Just as for  $\nu_2$ , the totally symmetric  $\nu_1$  stretching mode is, in first approximation, only Raman-active. Contrary to the  $\text{CH}_4$  case [43] and to what happens with  $\nu_2$  (see previous Section), rovibrational interaction-induced  $\nu_1$  lines are too weak to be observed in infrared absorption. Thus, up to now, high-resolution stimulated Raman spectroscopy was the only data source for this mode [24]. The Raman spectrum of  $\nu_1$  only displays a narrow  $Q$  branch for the  $\nu_1$  fundamental band, which strongly limits the number of assignable lines.

305

Recently, in the case of the  $\text{SF}_6$  molecule, we have shown another way to access such modes using the difference bands [44]. As a matter of fact, the far infrared region of  $\text{CF}_4$  near  $280\text{ cm}^{-1}$  displays a very nice  $\nu_1 - \nu_4$  band with a well-defined  $PQR$  branches structure. As described in Section 2, the

310

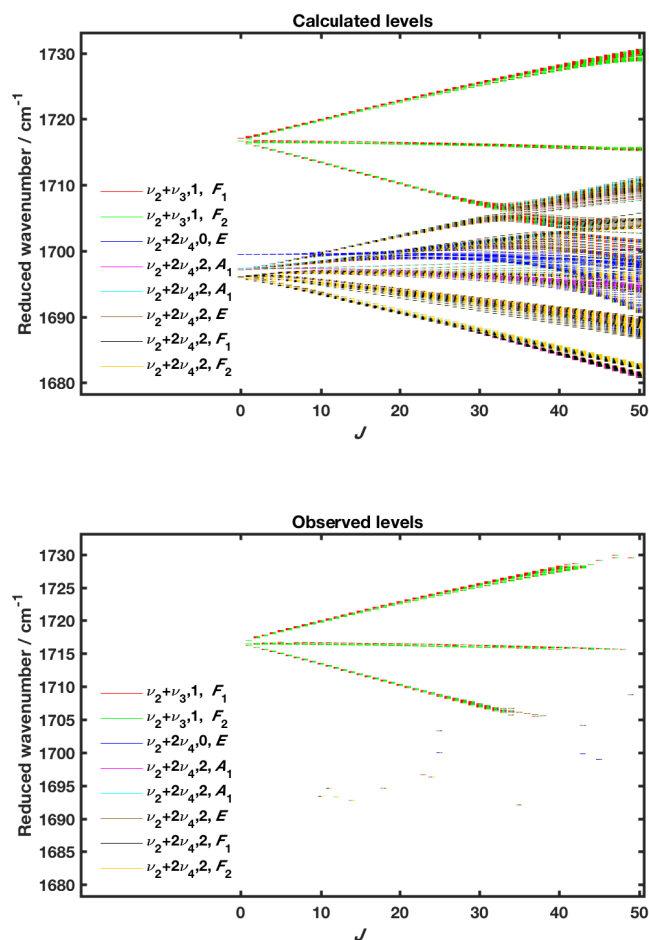


Figure 8: Observed and calculated reduced rovibrational energy levels for the  $\nu_2 + (\nu_3/2\nu_4)$  region, compared to the simulation. “Observed” levels correspond to those reached by assigned transitions. The colors represent the mixings between the different sublevels.

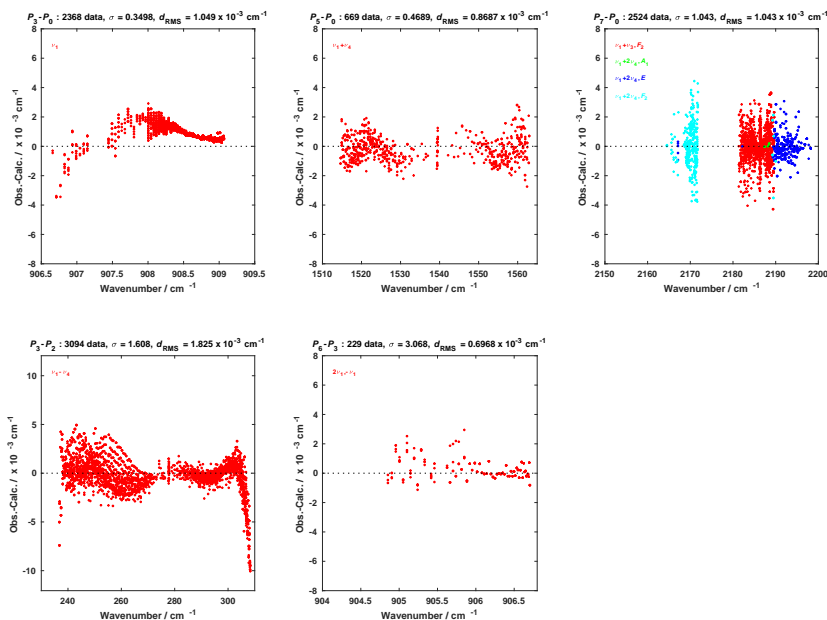


Figure 9: Fit residuals for line positions, in the case of the “ $\nu_1$ ” fit. In each spectral region, the colors correspond to the different vibrational sublevels.

synchrotron radiation was of a great help to obtain spectra with an excellent signal-to-noise ratio in this low wavenumber region.

315 It made it possible to perform a new global analysis including vibrational levels implying the  $\nu_1$  mode. This includes  $\nu_1$  and  $2\nu_1 - \nu_1$  Raman lines, and new spectra for  $\nu_1 - \nu_4$ ,  $\nu_1 + \nu_4$  and  $\nu_1 + (\nu_3/2\nu_4)$ . As explained in Section 3.2 and Table 4, this is referred as the “ $\nu_1$ ” scheme and fit. Parameters for the ground state,  $v_4 = 1$ ,  $v_4 = 2$  and  $v_3 = 1$  levels were fixed to the values obtained  
 320 in the “ $\nu_2$ ” fit. The reason is that, in this second case, we have less data and less parameters to fit. The line assignment and fit and simulation procedures using the XTDS and SPVIEW softwares [42] were totally similar to what we did for the “ $\nu_2$ ” fit (see previous Section).

Table 5 summarizes the fit statistics and Figure 9 displays the fit residuals  
 325 for line positions. Table B.7 in the Appendix lists the effective Hamiltonian parameter values.



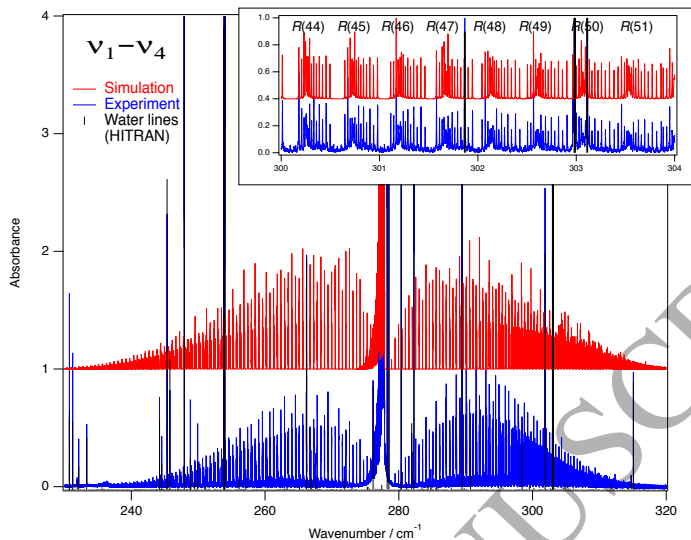


Figure 10: Measured  $\nu_1 - \nu_4$  difference band at 296 K and 18 Torr, compared to the simulation. Strong lines peaks are water vapor lines. The insert details a part of the  $R$  branch.

Figure 10 displays the  $\nu_1 - \nu_4$  region, compared to the simulation. The insert nicely illustrates the well-defined fine structure in the  $P$  branch. Figure 11 shows the  $\nu_1 + (\nu_3/2\nu_4)$  region. This one has 4 vibrational sublevels. Figure 12 displays the calculated and observed reduced energy levels, as defined by Equation 14.

#### 4.3. Discussion

In Ref. [24], we calculated the equilibrium bond length  $r_e$  of  $\text{CF}_4$ . This calculation is possible if one knows the value of the rotational constant  $B_0$  in the ground state and rotational constant differences in all the vibrational

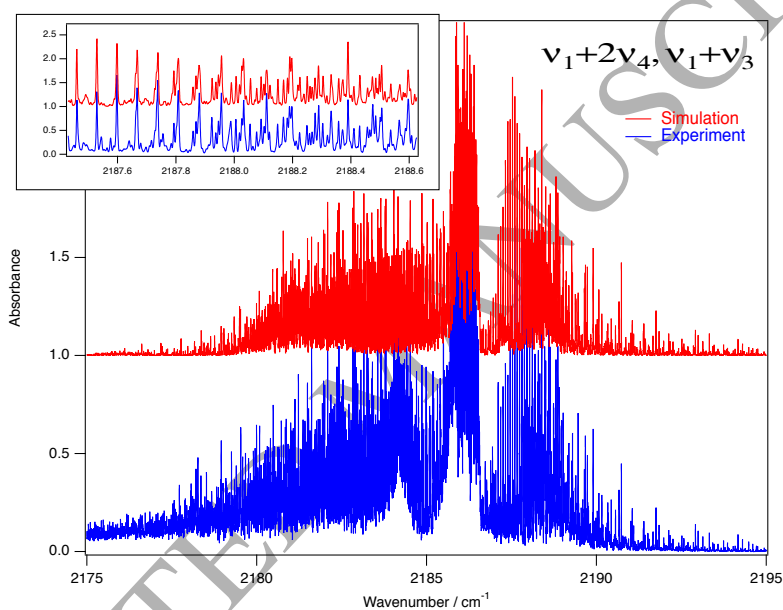


Figure 11: Measured spectrum of the  $\nu_1 + (\nu_3/2\nu_4)$  region at 295 K and 0.5 Torr, compared to the simulation. The insert details a part of the *R* branch. The strong *Q* branch near 2183  $\text{cm}^{-1}$  is the  $\nu_1 + \nu_3 + \nu_2 - \nu_2$  hot band, which is not simulated here.

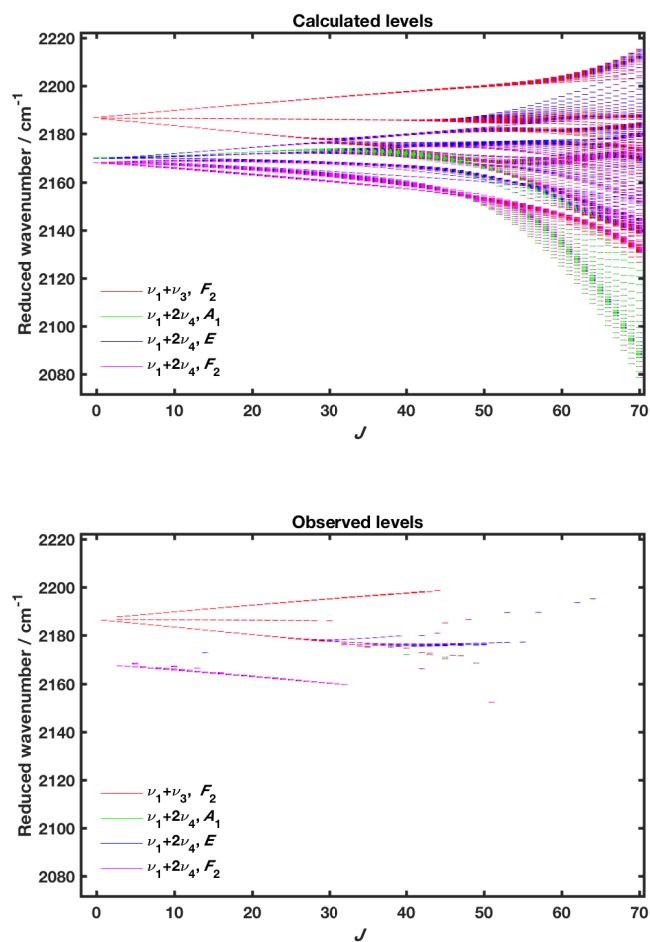


Figure 12: Observed and calculated reduced rovibrational energy levels for the  $\nu_1 + (\nu_3/2\nu_4)$  region, compared to the simulation. “Observed” levels correspond to those reached by assigned transitions. The colors represent the mixings between the different sublevels.

340 fundamental levels, say  $\Delta B_i = B_i - B_0$  ( $i = 1, \dots, 4$ ), where  $B_i$  is the rotational constant in an excited state with  $v_i = 1$ . The formula giving  $B_e$  is [36]:

$$B_e = B_0 - \frac{1}{2} \sum_{i=1}^4 d_i \Delta B_i, \quad (15)$$

where  $d_i$  is the normal mode degeneracy ( $d_1 = 1$ ,  $d_2 = 2$ ,  $d_3 = d_4 = 3$ ). The equilibrium bond length is:

$$r_e = \sqrt{\frac{3h}{64\pi^2 c m_F B_e}}, \quad (16)$$

$h$  being Planck's constant and  $m_F$  the fluorine atom's mass. The fitted parameters  $\Delta B_i = t_{\{i\}\{i\}}^{2(0,0A_1)\Gamma_v\Gamma_v}$  result from the two previous global fits. Using their values and standard deviation taken from Tables A.6 and B.7, we get:

$$B_e = 0.1924649(61) \text{ cm}^{-1}, \quad (17)$$

and thus:

$$r_e = 1.314860(21) \text{ \AA}. \quad (18)$$

This value is slightly smaller than in Ref. [24], but more precise.

## 5. Line intensities

350 Using the present results about the  $v_2 = v_3 = 1/v_4 = 2$  level ( $\nu_2 + (\nu_3/2\nu_4)$  dyad region) and of  $v_1 = v_3 = 1/v_4 = 2$  level ( $\nu_1 + (\nu_3/2\nu_4)$  dyad region), it becomes possible to calculate the  $\nu_3 + \nu_2 - \nu_2$ ,  $2\nu_4 + \nu_2 - \nu_2$ ,  $\nu_3 + \nu_1 - \nu_1$  and  $2\nu_3 + \nu_1 - \nu_1$  hot bands, in the  $\nu_3/2\nu_4$  region. In order to calculate absolute line intensities, we can use the  $\nu_3$  and  $2\nu_4$  dipole moment derivative parameters from  
 355 Ref. [45] (see also Section 5 of Ref. [23] for more explanations). These two values are sufficient and common to both cold and hot bands, when limiting the dipole moment operator expansion to degree 1. No further dipole moment refinement is possible, since no individual line intensity measurement is available, due to the spectroscopic congestion for this molecules, for which almost all observed  
 360 lines result from the overlap of several transitions.

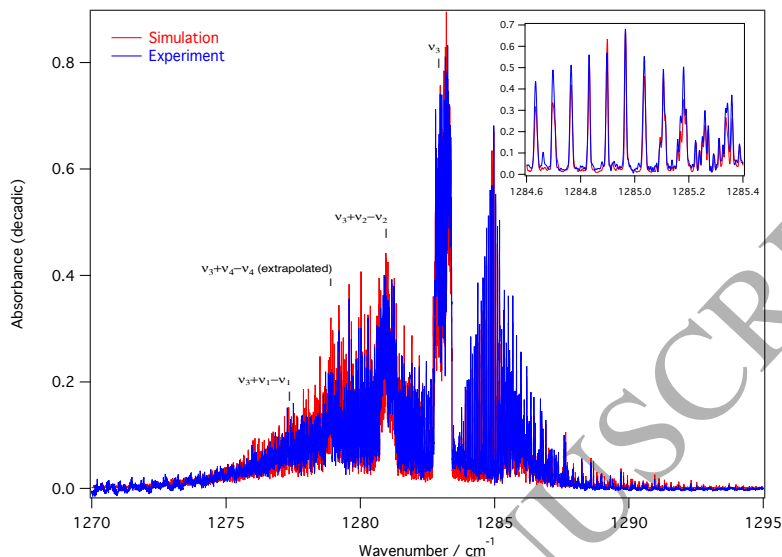


Figure 13: Experimental spectrum in the  $\nu_3$  region (296 K, 17 Torr), including hot bands, compared to the simulation. The insert details a part of the  $R$  branch. Most of the absorption intensity is well reproduced by the model.

It is even possible to extrapolate the  $\nu_3 + \nu_4 - \nu_4$  and  $3\nu_4 - \nu_4$  hot bands, although the corresponding vibrational levels could not be assigned in the current work. These would be a part of the polyad number 9 in the “ $\nu_2$ ” polyad scheme, whose effective Hamiltonian can be expanded up to order two only, leading to purely vibrational parameters only. We just set up the four parameters of the  $\nu_3 = \nu_4 = 1$  vibrational level (corresponding to the anharmonicity coefficient of its four sublevels) to a constant value of  $-4.7 \text{ cm}^{-1}$ , corresponding to a correct positioning of the  $\nu_3 + \nu_4 - \nu_4$   $Q$  branch (see Figures 13 and 14 below).

The resulting simulation at high resolution is shown on Figure 13, compared to the corresponding spectrum recorded in the present work. Figure 14 shows a low-resolution simulation, compared to a band profile from the PNNL database [46].

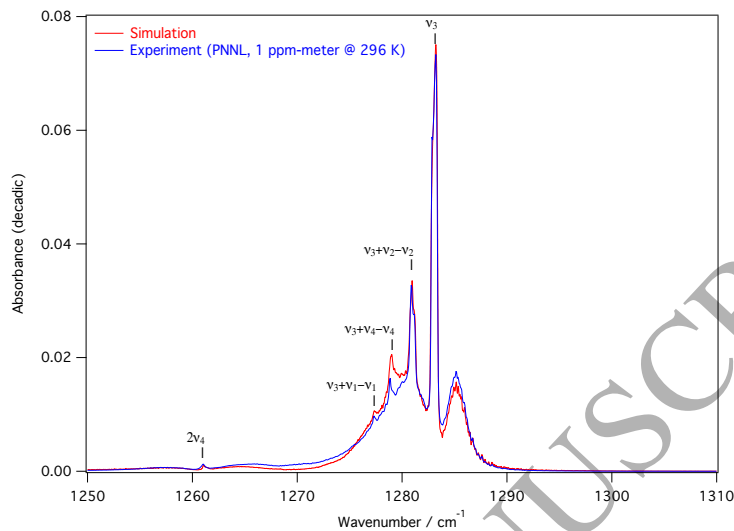


Figure 14: PNNL (see text and Ref. [46]) spectrum of the  $\nu_3$  region, including hot bands, at low resolution, compared to the simulation.

375 It appears clearly that we can now simulate a significant part (almost 92 %, see Figure 1) of the  $\nu_3$  region absorption intensity at room temperature, compared to the sole  $\nu_3$  fundamental simulation (which represents only 65.8 % of the intensity).

## 6. HITRAN and TFMcCaSDa updates

380 Although the present work was not fully ready on time for the recent HITRAN 2016 [18] and GEISA 2015 [20] updates, a new partial calculated line list including  $\nu_3$  and  $\nu_3 + \nu_2 - \nu_2$  only could be used for the HITRAN 2016 update itself. Then, the complete  $\text{CF}_4$  line list in the  $\nu_3$  region, including hot bands based on  $\nu_1$ ,  $\nu_2$  and  $\nu_4$  (see previous Section) has now been accepted by these  
385 databases for an intermediate online update.

Moreover, we have previously setup a database named TFMcCaSDa (*Tetra-Fluoro-Methane Calculated Spectroscopic Database*) of the calculated  $\text{CF}_4$  spectroscopic lines in the framework of the *Virtual Atomic and Molecular Data Cen-*

390 *tre* (VAMDC) [47, 48, 49]. This one has also been updated using the present data. It is accessible through the VAMDC portal (<http://portal.vamdc.org>).

## 7. Conclusion

The two global fits presented in this study constitute the most extensive study of  $\text{CF}_4$  rovibrational levels to date. This enables a significant update of the public spectroscopic databases concerning this molecule. Our results should 395 allow a much better modeling of the atmospheric absorption and improved concentration retrievals from satellite, balloon-borne and ground-based data.

The next step will be to include even more bands in the analysis, possibly in a unique global polyad scheme, if possible. In particular, the  $\nu_3 + \nu_4$  band should be assigned and analyzed in detail. The knowledge of this one would 400 lead to the calculation of the  $\nu_3 + \nu_4 - \nu_4$  hot band, which is the second one in terms of absorption intensity, representing 9.2 % of it in the  $\nu_3$  region at room temperature (here, we could just extrapolate this band). This kind of rovibrational system is quite challenging to analyze, since the  $v_3 = v_4 = 1$  vibrational level has four sublevels whose exact position is still unknown. Presumably, *ab* 405 *initio* calculations would be of help to get initial values for these anharmonicity coefficients. Moreover, this band should interact with  $3\nu_4$ .

Another important challenge would be to accurately measure the line intensities and to determine the experimental dipole moment parameters, instead of using previously calculated values. This is not an easy task, for such a heavy 410 molecule with a very congested spectrum consisting mostly of many overlapping lines.

## 8. Acknowledgements

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### Appendix A. Effective Hamiltonian parameters from the “ $\nu_2$ ” fit.

Table A.6 below gives the fitted values of the effective Hamiltonian parameters of  $\text{CF}_4$  for the “ $\nu_2$ ” fit (see Section 4.1). Notations correspond to Equation (3). Standard deviation is given in parentheses in the unit of the last two digits.

Parameter	Polyad	Order	$\Omega(K, nC)$	$\{s\} C_v$	$\{s'\} C'_v$	Value / $\text{cm}^{-1}$
1	GS	0	2(0, 0A <sub>1</sub> )	0000A <sub>1</sub>	0000A <sub>1</sub>	1.9119582(35) $\times 10^{-1}$
2	GS	2	4(0, 0A <sub>1</sub> )	0000A <sub>1</sub>	0000A <sub>1</sub>	-6.380(33) $\times 10^{-8}$
3	GS	2	4(4, 0A <sub>1</sub> )	0000A <sub>1</sub>	0000A <sub>1</sub>	-3.433(17) $\times 10^{-9}$
4	GS	4	6(0, 0A <sub>1</sub> )	0000A <sub>1</sub>	0000A <sub>1</sub>	2.30(10) $\times 10^{-12}$
5	GS	4	6(4, 0A <sub>1</sub> )	0000A <sub>1</sub>	0000A <sub>1</sub>	-1.902(44) $\times 10^{-13}$
6	GS	4	6(6, 0A <sub>1</sub> )	0000A <sub>1</sub>	0000A <sub>1</sub>	-1.010(12) $\times 10^{-13}$
7	GS	6	8(0, 0A <sub>1</sub> )	0000A <sub>1</sub>	0000A <sub>1</sub>	8.544(92) $\times 10^{-16}$
8	GS	6	8(4, 0A <sub>1</sub> )	0000A <sub>1</sub>	0000A <sub>1</sub>	4.654(30) $\times 10^{-17}$
9	GS	6	8(6, 0A <sub>1</sub> )	0000A <sub>1</sub>	0000A <sub>1</sub>	-1.054(11) $\times 10^{-17}$
10	GS	6	8(8, 0A <sub>1</sub> )	0000A <sub>1</sub>	0000A <sub>1</sub>	4.73(45) $\times 10^{-19}$
16	$\nu_2$	0	0(0, 0A <sub>1</sub> )	0100E	0100E	435.380220(40)
17	$\nu_2$	2	2(0, 0A <sub>1</sub> )	0100E	0100E	-3.37820(67) $\times 10^{-4}$
18	$\nu_2$	2	2(2, 0E)	0100E	0100E	-1.04544(70) $\times 10^{-4}$
19	$\nu_2$	3	3(3, 0A <sub>2</sub> )	0100E	0100E	2.9171(54) $\times 10^{-7}$
20	$\nu_2$	4	4(0, 0A <sub>1</sub> )	0100E	0100E	7.96(32) $\times 10^{-10}$
21	$\nu_2$	4	4(2, 0E)	0100E	0100E	1.440(26) $\times 10^{-9}$
22	$\nu_2$	4	4(4, 0A <sub>1</sub> )	0100E	0100E	2.681(45) $\times 10^{-10}$
23	$\nu_2$	4	4(4, 0E)	0100E	0100E	-1.95(14) $\times 10^{-10}$
24	$\nu_2$	5	5(3, 0A <sub>2</sub> )	0100E	0100E	-2.270(77) $\times 10^{-12}$
25	$\nu_2$	6	6(0, 0A <sub>1</sub> )	0100E	0100E	-7.66(36) $\times 10^{-14}$
26	$\nu_2$	6	6(2, 0E)	0100E	0100E	1.243(19) $\times 10^{-13}$
28	$\nu_2$	6	6(4, 0E)	0100E	0100E	-3.50(14) $\times 10^{-14}$
42	$\nu_4$	0	0(0, 0A <sub>1</sub> )	0001F <sub>2</sub>	0001F <sub>2</sub>	631.059781(74)
43	$\nu_4$	1	1(1, 0F <sub>1</sub> )	0001F <sub>2</sub>	0001F <sub>2</sub>	-2.947805(68) $\times 10^{-1}$
44	$\nu_4$	2	2(0, 0A <sub>1</sub> )	0001F <sub>2</sub>	0001F <sub>2</sub>	1.0277(17) $\times 10^{-4}$
45	$\nu_4$	2	2(2, 0E)	0001F <sub>2</sub>	0001F <sub>2</sub>	-8.569(17) $\times 10^{-5}$
46	$\nu_4$	2	2(2, 0F <sub>2</sub> )	0001F <sub>2</sub>	0001F <sub>2</sub>	1.7218(35) $\times 10^{-4}$
47	$\nu_4$	3	3(1, 0F <sub>1</sub> )	0001F <sub>2</sub>	0001F <sub>2</sub>	-3.327(35) $\times 10^{-7}$
48	$\nu_4$	3	3(3, 0F <sub>1</sub> )	0001F <sub>2</sub>	0001F <sub>2</sub>	-5.081(15) $\times 10^{-7}$
49	$\nu_4$	4	4(0, 0A <sub>1</sub> )	0001F <sub>2</sub>	0001F <sub>2</sub>	-2.01(20) $\times 10^{-9}$

Table A.6: Effective Hamiltonian parameters.



Table A.6 (continued).

Parameter	Polyad	Order	$\Omega(K, nC)$	$\{s\} C_v$	$\{s'\} C'_v$	Value / $\text{cm}^{-1}$
50	$\nu_4$	4	4(2, 0E)	0001F <sub>2</sub>	0001F <sub>2</sub>	-3.84(23) $\times 10^{-9}$
51	$\nu_4$	4	4(2, 0F <sub>2</sub> )	0001F <sub>2</sub>	0001F <sub>2</sub>	-3.11(26) $\times 10^{-9}$
52	$\nu_4$	4	4(4, 0A <sub>1</sub> )	0001F <sub>2</sub>	0001F <sub>2</sub>	1.177(63) $\times 10^{-9}$
53	$\nu_4$	4	4(4, 0E)	0001F <sub>2</sub>	0001F <sub>2</sub>	6.03(36) $\times 10^{-9}$
54	$\nu_4$	4	4(4, 0F <sub>2</sub> )	0001F <sub>2</sub>	0001F <sub>2</sub>	4.73(26) $\times 10^{-9}$
55	$\nu_4$	5	5(1, 0F <sub>1</sub> )	0001F <sub>2</sub>	0001F <sub>2</sub>	1.696(42) $\times 10^{-11}$
56	$\nu_4$	5	5(3, 0F <sub>1</sub> )	0001F <sub>2</sub>	0001F <sub>2</sub>	-2.556(75) $\times 10^{-11}$
57	$\nu_4$	5	5(5, 0F <sub>1</sub> )	0001F <sub>2</sub>	0001F <sub>2</sub>	1.106(81) $\times 10^{-11}$
59	$\nu_4$	6	6(0, 0A <sub>1</sub> )	0001F <sub>2</sub>	0001F <sub>2</sub>	-4.84(16) $\times 10^{-13}$
90	$2\nu_2$	2	0(0, 0A <sub>1</sub> )	0200A <sub>1</sub>	0200A <sub>1</sub>	-2.841606(78)
91	$2\nu_2$	4	2(0, 0A <sub>1</sub> )	0200A <sub>1</sub>	0200A <sub>1</sub>	2.5371(78) $\times 10^{-5}$
97	$2\nu_2$	4	2(2, 0E)	0200A <sub>1</sub>	0200E	-2.19(32) $\times 10^{-6}$
98	$2\nu_2$	6	4(2, 0E)	0200A <sub>1</sub>	0200E	-2.32(28) $\times 10^{-10}$
104	$2\nu_2$	2	0(0, 0A <sub>1</sub> )	0200E	0200E	3.8436(17) $\times 10^{-1}$
105	$2\nu_2$	4	2(0, 0A <sub>1</sub> )	0200E	0200E	1.84(10) $\times 10^{-6}$
119	$\nu_2 + \nu_4$	2	0(0, 0A <sub>1</sub> )	0101F <sub>1</sub>	0101F <sub>1</sub>	2.6540(31) $\times 10^{-1}$
120	$\nu_2 + \nu_4$	3	1(1, 0F <sub>1</sub> )	0101F <sub>1</sub>	0101F <sub>1</sub>	5.34(17) $\times 10^{-4}$
121	$\nu_2 + \nu_4$	4	2(0, 0A <sub>1</sub> )	0101F <sub>1</sub>	0101F <sub>1</sub>	9.0(1.2) $\times 10^{-6}$
146	$\nu_2 + \nu_4$	3	1(1, 0F <sub>1</sub> )	0101F <sub>1</sub>	0101F <sub>2</sub>	-1.754(12) $\times 10^{-3}$
147	$\nu_2 + \nu_4$	4	2(2, 0E)	0101F <sub>1</sub>	0101F <sub>2</sub>	-7.85(69) $\times 10^{-6}$
148	$\nu_2 + \nu_4$	4	2(2, 0F <sub>2</sub> )	0101F <sub>1</sub>	0101F <sub>2</sub>	-4.41(41) $\times 10^{-6}$
176	$\nu_2 + \nu_4$	2	0(0, 0A <sub>1</sub> )	0101F <sub>2</sub>	0101F <sub>2</sub>	-3.2221(18) $\times 10^{-1}$
177	$\nu_2 + \nu_4$	3	1(1, 0F <sub>1</sub> )	0101F <sub>2</sub>	0101F <sub>2</sub>	-3.774(17) $\times 10^{-3}$
178	$\nu_2 + \nu_4$	4	2(0, 0A <sub>1</sub> )	0101F <sub>2</sub>	0101F <sub>2</sub>	-1.81(12) $\times 10^{-5}$
203	$\nu_3/2\nu_4$	0	0(0, 0A <sub>1</sub> )	0010F <sub>2</sub>	0010F <sub>2</sub>	1283.236(19)
204	$\nu_3/2\nu_4$	1	1(1, 0F <sub>1</sub> )	0010F <sub>2</sub>	0010F <sub>2</sub>	6.4617(72) $\times 10^{-1}$
205	$\nu_3/2\nu_4$	2	2(0, 0A <sub>1</sub> )	0010F <sub>2</sub>	0010F <sub>2</sub>	-6.079(36) $\times 10^{-4}$
206	$\nu_3/2\nu_4$	2	2(2, 0E)	0010F <sub>2</sub>	0010F <sub>2</sub>	1.244(13) $\times 10^{-4}$
207	$\nu_3/2\nu_4$	2	2(2, 0F <sub>2</sub> )	0010F <sub>2</sub>	0010F <sub>2</sub>	-2.7521(88) $\times 10^{-4}$
208	$\nu_3/2\nu_4$	3	3(1, 0F <sub>1</sub> )	0010F <sub>2</sub>	0010F <sub>2</sub>	1.787(86) $\times 10^{-6}$

Table A.6 (continued).

Parameter	Polyad	Order	$\Omega(K, nC)$	$\{s\}$ $C_v$	$\{s'\}$ $C'_v$	Value / $\text{cm}^{-1}$	
209	$\nu_3/2\nu_4$	3	3(3, 0F <sub>1</sub> )	0010F <sub>2</sub>	0010F <sub>2</sub>	9.88(97)	$\times 10^{-8}$
211	$\nu_3/2\nu_4$	4	4(2, 0E)	0010F <sub>2</sub>	0010F <sub>2</sub>	7.8(1.4)	$\times 10^{-9}$
212	$\nu_3/2\nu_4$	4	4(2, 0F <sub>2</sub> )	0010F <sub>2</sub>	0010F <sub>2</sub>	-1.54(15)	$\times 10^{-8}$
213	$\nu_3/2\nu_4$	4	4(4, 0A <sub>1</sub> )	0010F <sub>2</sub>	0010F <sub>2</sub>	7.82(69)	$\times 10^{-10}$
214	$\nu_3/2\nu_4$	4	4(4, 0E)	0010F <sub>2</sub>	0010F <sub>2</sub>	-1.19(23)	$\times 10^{-8}$
215	$\nu_3/2\nu_4$	4	4(4, 0F <sub>2</sub> )	0010F <sub>2</sub>	0010F <sub>2</sub>	-1.11(16)	$\times 10^{-8}$
220	$\nu_3/2\nu_4$	6	6(0, 0A <sub>1</sub> )	0010F <sub>2</sub>	0010F <sub>2</sub>	-4.51(12)	$\times 10^{-12}$
251	$\nu_3/2\nu_4$	3	2(2, 0F <sub>2</sub> )	0010F <sub>2</sub>	0002A <sub>1</sub>	1.65(12)	$\times 10^{-4}$
252	$\nu_3/2\nu_4$	4	3(3, 0F <sub>2</sub> )	0010F <sub>2</sub>	0002A <sub>1</sub>	-5.43(21)	$\times 10^{-6}$
265	$\nu_3/2\nu_4$	2	1(1, 0F <sub>1</sub> )	0010F <sub>2</sub>	0002E	-2.19(10)	$\times 10^{-2}$
266	$\nu_3/2\nu_4$	3	2(2, 0F <sub>2</sub> )	0010F <sub>2</sub>	0002E	9.30(16)	$\times 10^{-4}$
295	$\nu_3/2\nu_4$	1	0(0, 0A <sub>1</sub> )	0010F <sub>2</sub>	0002F <sub>2</sub>	3.326(63)	
296	$\nu_3/2\nu_4$	2	1(1, 0F <sub>1</sub> )	0010F <sub>2</sub>	0002F <sub>2</sub>	5.05(25)	$\times 10^{-2}$
297	$\nu_3/2\nu_4$	3	2(0, 0A <sub>1</sub> )	0010F <sub>2</sub>	0002F <sub>2</sub>	6.56(12)	$\times 10^{-4}$
341	$\nu_3/2\nu_4$	2	0(0, 0A <sub>1</sub> )	0002A <sub>1</sub>	0002A <sub>1</sub>	-2.65(19)	$\times 10^{-1}$
342	$\nu_3/2\nu_4$	4	2(0, 0A <sub>1</sub> )	0002A <sub>1</sub>	0002A <sub>1</sub>	2.86(57)	$\times 10^{-4}$
343	$\nu_3/2\nu_4$	6	4(0, 0A <sub>1</sub> )	0002A <sub>1</sub>	0002A <sub>1</sub>	-3.86(23)	$\times 10^{-7}$
344	$\nu_3/2\nu_4$	6	4(4, 0A <sub>1</sub> )	0002A <sub>1</sub>	0002A <sub>1</sub>	2.59(30)	$\times 10^{-8}$
348	$\nu_3/2\nu_4$	4	2(2, 0E)	0002A <sub>1</sub>	0002E	2.640(40)	$\times 10^{-4}$
350	$\nu_3/2\nu_4$	6	4(4, 0E)	0002A <sub>1</sub>	0002E	5.76(27)	$\times 10^{-8}$
355	$\nu_3/2\nu_4$	4	2(2, 0F <sub>2</sub> )	0002A <sub>1</sub>	0002F <sub>2</sub>	-4.049(82)	$\times 10^{-4}$
356	$\nu_3/2\nu_4$	5	3(3, 0F <sub>2</sub> )	0002A <sub>1</sub>	0002F <sub>2</sub>	-6.75(42)	$\times 10^{-6}$
358	$\nu_3/2\nu_4$	6	4(4, 0F <sub>2</sub> )	0002A <sub>1</sub>	0002F <sub>2</sub>	1.555(42)	$\times 10^{-7}$
365	$\nu_3/2\nu_4$	2	0(0, 0A <sub>1</sub> )	0002E	0002E	-9.12(27)	$\times 10^{-3}$
366	$\nu_3/2\nu_4$	4	2(0, 0A <sub>1</sub> )	0002E	0002E	1.387(13)	$\times 10^{-3}$
367	$\nu_3/2\nu_4$	4	2(2, 0E)	0002E	0002E	1.203(12)	$\times 10^{-3}$
380	$\nu_3/2\nu_4$	3	1(1, 0F <sub>1</sub> )	0002E	0002F <sub>2</sub>	1.525(20)	$\times 10^{-2}$
381	$\nu_3/2\nu_4$	4	2(2, 0F <sub>2</sub> )	0002E	0002F <sub>2</sub>	-7.590(84)	$\times 10^{-4}$
382	$\nu_3/2\nu_4$	5	3(1, 0F <sub>1</sub> )	0002E	0002F <sub>2</sub>	-8.20(16)	$\times 10^{-6}$
383	$\nu_3/2\nu_4$	5	3(3, 0F <sub>1</sub> )	0002E	0002F <sub>2</sub>	1.286(56)	$\times 10^{-6}$

Table A.6 (continued).

Parameter	Polyad	Order	$\Omega(K, nC)$	$\{s\} C_v$	$\{s'\} C'_v$	Value / $\text{cm}^{-1}$
384	$\nu_3/2\nu_4$	5	3(3, 0F <sub>2</sub> )	0002E	0002F <sub>2</sub>	-2.656(59) $\times 10^{-6}$
385	$\nu_3/2\nu_4$	6	4(2, 0F <sub>2</sub> )	0002E	0002F <sub>2</sub>	-3.67(16) $\times 10^{-8}$
386	$\nu_3/2\nu_4$	6	4(4, 0F <sub>1</sub> )	0002E	0002F <sub>2</sub>	-1.056(71) $\times 10^{-8}$
387	$\nu_3/2\nu_4$	6	4(4, 0F <sub>2</sub> )	0002E	0002F <sub>2</sub>	-5.32(14) $\times 10^{-8}$
400	$\nu_3/2\nu_4$	2	0(0, 0A <sub>1</sub> )	0002F <sub>2</sub>	0002F <sub>2</sub>	-1.204(19)
401	$\nu_3/2\nu_4$	3	1(1, 0F <sub>1</sub> )	0002F <sub>2</sub>	0002F <sub>2</sub>	3.999(72) $\times 10^{-2}$
402	$\nu_3/2\nu_4$	4	2(0, 0A <sub>1</sub> )	0002F <sub>2</sub>	0002F <sub>2</sub>	-8.642(97) $\times 10^{-4}$
403	$\nu_3/2\nu_4$	4	2(2, 0E)	0002F <sub>2</sub>	0002F <sub>2</sub>	1.374(14) $\times 10^{-3}$
404	$\nu_3/2\nu_4$	4	2(2, 0F <sub>2</sub> )	0002F <sub>2</sub>	0002F <sub>2</sub>	-1.1851(98) $\times 10^{-3}$
405	$\nu_3/2\nu_4$	5	3(1, 0F <sub>1</sub> )	0002F <sub>2</sub>	0002F <sub>2</sub>	-1.022(91) $\times 10^{-6}$
408	$\nu_3/2\nu_4$	6	4(2, 0E)	0002F <sub>2</sub>	0002F <sub>2</sub>	3.41(13) $\times 10^{-8}$
409	$\nu_3/2\nu_4$	6	4(2, 0F <sub>2</sub> )	0002F <sub>2</sub>	0002F <sub>2</sub>	-2.72(13) $\times 10^{-8}$
410	$\nu_3/2\nu_4$	6	4(4, 0A <sub>1</sub> )	0002F <sub>2</sub>	0002F <sub>2</sub>	-5.56(44) $\times 10^{-9}$
645 411	$\nu_3/2\nu_4$	6	4(4, 0E)	0002F <sub>2</sub>	0002F <sub>2</sub>	-4.64(16) $\times 10^{-8}$
412	$\nu_3/2\nu_4$	6	4(4, 0F <sub>2</sub> )	0002F <sub>2</sub>	0002F <sub>2</sub>	-2.72(11) $\times 10^{-8}$
505	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	2	0(0, 0A <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	-2.0761(23)
506	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	3	1(1, 0F <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	1.0433(72) $\times 10^{-2}$
507	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	2(0, 0A <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	-4.72(14) $\times 10^{-4}$
508	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	2(2, 0E)	0110F <sub>1</sub>	0110F <sub>1</sub>	-3.71(11) $\times 10^{-4}$
509	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	2(2, 0F <sub>2</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	5.21(17) $\times 10^{-4}$
510	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	3(1, 0F <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	3.8(1.0) $\times 10^{-7}$
511	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	3(3, 0F <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	6.15(55) $\times 10^{-6}$
512	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	4(0, 0A <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	-5.359(90) $\times 10^{-7}$
513	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	4(2, 0E)	0110F <sub>1</sub>	0110F <sub>1</sub>	4.667(62) $\times 10^{-7}$
514	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	4(2, 0F <sub>2</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	-6.523(63) $\times 10^{-7}$
515	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	4(4, 0A <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	-5.80(26) $\times 10^{-8}$
516	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	4(4, 0E)	0110F <sub>1</sub>	0110F <sub>1</sub>	-3.701(59) $\times 10^{-7}$
517	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	4(4, 0F <sub>2</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	6.516(99) $\times 10^{-7}$
518	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	5(1, 0F <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	1.974(85) $\times 10^{-9}$
519	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	5(3, 0F <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	-2.595(91) $\times 10^{-9}$

Table A.6 (continued).

Parameter	Polyad	Order	$\Omega(K, nC)$	$\{s\}$ $C_v$	$\{s'\}$ $C'_v$	Value / $\text{cm}^{-1}$
520	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	5(5, 0F <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	6.35(49) $\times 10^{-10}$
521	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	5(5, 1F <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	3.08(11) $\times 10^{-9}$
522	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	6(0, 0A <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	1.068(34) $\times 10^{-10}$
523	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	6(2, 0E )	0110F <sub>1</sub>	0110F <sub>1</sub>	5.500(81) $\times 10^{-11}$
524	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	6(2, 0F <sub>2</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	-2.47(10) $\times 10^{-11}$
525	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	6(4, 0A <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	-9.42(26) $\times 10^{-12}$
526	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	6(4, 0E )	0110F <sub>1</sub>	0110F <sub>1</sub>	-5.76(11) $\times 10^{-11}$
527	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	6(4, 0F <sub>2</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	9.79(18) $\times 10^{-11}$
528	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	6(6, 0A <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	-1.16(12) $\times 10^{-12}$
529	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	6(6, 0E )	0110F <sub>1</sub>	0110F <sub>1</sub>	3.32(32) $\times 10^{-12}$
530	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	6(6, 0F <sub>2</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	-5.07(43) $\times 10^{-12}$
531	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	6(6, 1F <sub>2</sub> )	0110F <sub>1</sub>	0110F <sub>1</sub>	-5.90(61) $\times 10^{-12}$
532	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	3	1(1, 0F <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>2</sub>	2.306(98) $\times 10^{-3}$
533	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	2(2, 0E )	0110F <sub>1</sub>	0110F <sub>2</sub>	3.48(11) $\times 10^{-4}$
534	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	2(2, 0F <sub>2</sub> )	0110F <sub>1</sub>	0110F <sub>2</sub>	2.22(16) $\times 10^{-5}$
535	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	3(1, 0F <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>2</sub>	1.62(17) $\times 10^{-6}$
537	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	3(3, 0F <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>2</sub>	-1.67(17) $\times 10^{-6}$
538	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	3(3, 0F <sub>2</sub> )	0110F <sub>1</sub>	0110F <sub>2</sub>	1.08(21) $\times 10^{-6}$
539	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	4(2, 0E )	0110F <sub>1</sub>	0110F <sub>2</sub>	-3.556(46) $\times 10^{-7}$
540	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	4(2, 0F <sub>2</sub> )	0110F <sub>1</sub>	0110F <sub>2</sub>	1.20(42) $\times 10^{-8}$
541	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	4(4, 0E )	0110F <sub>1</sub>	0110F <sub>2</sub>	2.138(30) $\times 10^{-7}$
542	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	4(4, 0F <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>2</sub>	3.193(62) $\times 10^{-7}$
543	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	4(4, 0F <sub>2</sub> )	0110F <sub>1</sub>	0110F <sub>2</sub>	1.42(46) $\times 10^{-8}$
544	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	5(1, 0F <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>2</sub>	1.197(45) $\times 10^{-9}$
545	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	5(3, 0A <sub>2</sub> )	0110F <sub>1</sub>	0110F <sub>2</sub>	1.205(26) $\times 10^{-9}$
546	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	5(3, 0F <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>2</sub>	2.328(52) $\times 10^{-9}$
547	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	5(3, 0F <sub>2</sub> )	0110F <sub>1</sub>	0110F <sub>2</sub>	-1.168(51) $\times 10^{-9}$
548	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	5(5, 0E )	0110F <sub>1</sub>	0110F <sub>2</sub>	1.693(51) $\times 10^{-9}$
549	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	5(5, 0F <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>2</sub>	7.45(41) $\times 10^{-10}$
550	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	5(5, 1F <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>2</sub>	-1.530(44) $\times 10^{-9}$

Table A.6 (continued).

Parameter	Polyad	Order	$\Omega(K, nC)$	$\{s\}$ $C_v$	$\{s'\}$ $C'_v$	Value / $\text{cm}^{-1}$
551	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	5(5, 0F <sub>2</sub> )	0110F <sub>1</sub>	0110F <sub>2</sub>	1.490(55) $\times 10^{-9}$
552	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	6(2, 0E)	0110F <sub>1</sub>	0110F <sub>2</sub>	-3.729(66) $\times 10^{-11}$
553	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	6(2, 0F <sub>2</sub> )	0110F <sub>1</sub>	0110F <sub>2</sub>	1.506(95) $\times 10^{-11}$
555	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	6(4, 0F <sub>1</sub> )	0110F <sub>1</sub>	0110F <sub>2</sub>	1.374(41) $\times 10^{-11}$
556	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	6(4, 0F <sub>2</sub> )	0110F <sub>1</sub>	0110F <sub>2</sub>	-1.352(55) $\times 10^{-11}$
557	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	6(6, 0A <sub>2</sub> )	0110F <sub>1</sub>	0110F <sub>2</sub>	-2.65(22) $\times 10^{-12}$
562	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	2	0(0, 0A <sub>1</sub> )	0110F <sub>2</sub>	0110F <sub>2</sub>	-2.4859(24)
563	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	3	1(1, 0F <sub>1</sub> )	0110F <sub>2</sub>	0110F <sub>2</sub>	1.6860(93) $\times 10^{-2}$
564	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	2(0, 0A <sub>1</sub> )	0110F <sub>2</sub>	0110F <sub>2</sub>	4.90(14) $\times 10^{-4}$
565	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	2(2, 0E)	0110F <sub>2</sub>	0110F <sub>2</sub>	3.96(13) $\times 10^{-4}$
566	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	2(2, 0F <sub>2</sub> )	0110F <sub>2</sub>	0110F <sub>2</sub>	-6.87(17) $\times 10^{-4}$
567	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	3(1, 0F <sub>1</sub> )	0110F <sub>2</sub>	0110F <sub>2</sub>	-5.78(50) $\times 10^{-6}$
569	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	4(0, 0A <sub>1</sub> )	0110F <sub>2</sub>	0110F <sub>2</sub>	3.885(98) $\times 10^{-7}$
570	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	4(2, 0E)	0110F <sub>2</sub>	0110F <sub>2</sub>	-2.427(50) $\times 10^{-7}$
571	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	4(2, 0F <sub>2</sub> )	0110F <sub>2</sub>	0110F <sub>2</sub>	9.748(88) $\times 10^{-7}$
572	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	4(4, 0A <sub>1</sub> )	0110F <sub>2</sub>	0110F <sub>2</sub>	7.03(29) $\times 10^{-8}$
573	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	4(4, 0E)	0110F <sub>2</sub>	0110F <sub>2</sub>	2.86(49) $\times 10^{-8}$
574	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	4(4, 0F <sub>2</sub> )	0110F <sub>2</sub>	0110F <sub>2</sub>	-3.191(57) $\times 10^{-7}$
575	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	5(1, 0F <sub>1</sub> )	0110F <sub>2</sub>	0110F <sub>2</sub>	2.314(56) $\times 10^{-9}$
576	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	5(3, 0F <sub>1</sub> )	0110F <sub>2</sub>	0110F <sub>2</sub>	-9.64(51) $\times 10^{-10}$
577	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	5(5, 0F <sub>1</sub> )	0110F <sub>2</sub>	0110F <sub>2</sub>	1.295(35) $\times 10^{-9}$
578	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	5(5, 1F <sub>1</sub> )	0110F <sub>2</sub>	0110F <sub>2</sub>	8.04(50) $\times 10^{-10}$
579	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	6(0, 0A <sub>1</sub> )	0110F <sub>2</sub>	0110F <sub>2</sub>	-1.279(19) $\times 10^{-10}$
580	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	6(2, 0E)	0110F <sub>2</sub>	0110F <sub>2</sub>	-2.142(30) $\times 10^{-11}$
581	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	6(2, 0F <sub>2</sub> )	0110F <sub>2</sub>	0110F <sub>2</sub>	1.0103(91) $\times 10^{-10}$
603	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	1(1, 0F <sub>1</sub> )	0110F <sub>1</sub>	0102A <sub>1</sub>	2.63(11) $\times 10^{-2}$
611	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	2(2, 0F <sub>2</sub> )	0110F <sub>1</sub>	0102A <sub>2</sub>	-9.93(25) $\times 10^{-4}$
674	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	2(2, 0F <sub>2</sub> )	0110F <sub>2</sub>	0102E	-2.468(45) $\times 10^{-3}$
693	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	1(1, 0F <sub>1</sub> )	0110F <sub>2</sub>	0102A <sub>2</sub>	1.84(15) $\times 10^{-2}$
701	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	1(1, 0F <sub>1</sub> )	0110F <sub>2</sub>	0102E	3.48(11) $\times 10^{-2}$

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Table A.6 (continued).

Parameter	Polyad	Order	$\Omega(K, nC)$	$\{s\} C_v$	$\{s'\} C'_v$	Value / $\text{cm}^{-1}$
757	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	0(0, 0A <sub>1</sub> )	0102E	0102E	2.253(51)
771	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	0(0, 0A <sub>1</sub> )	0102E	0102E	3.79(48) $\times 10^{-1}$
779	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	1(1, 0F <sub>1</sub> )	0102E	0102F <sub>1</sub>	-5.64(32) $\times 10^{-2}$
787	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	1(1, 0F <sub>1</sub> )	0102E	0102F <sub>2</sub>	-8.03(28) $\times 10^{-2}$
795	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	0(0, 0A <sub>1</sub> )	0102A <sub>1</sub>	0102A <sub>1</sub>	4.61(23) $\times 10^{-1}$
796	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	2(0, 0A <sub>1</sub> )	0102A <sub>1</sub>	0102A <sub>1</sub>	-1.879(33) $\times 10^{-3}$
811	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	0(0, 0A <sub>1</sub> )	0102A <sub>2</sub>	0102A <sub>2</sub>	-6.78(32) $\times 10^{-1}$
815	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	2(2, 0E)	0102A <sub>2</sub>	0102E	-6.408(79) $\times 10^{-4}$
826	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	0(0, 0A <sub>1</sub> )	0102E	0102E	-2.913(72) $\times 10^{-1}$
834	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	1(1, 0F <sub>1</sub> )	0102E	0102F <sub>1</sub>	-3.026(62) $\times 10^{-2}$
842	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	1(1, 0F <sub>1</sub> )	0102E	0102F <sub>2</sub>	-1.296(78) $\times 10^{-2}$
850	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	0(0, 0A <sub>1</sub> )	0102F <sub>1</sub>	0102F <sub>1</sub>	6.579(60) $\times 10^{-1}$
851	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	1(1, 0F <sub>1</sub> )	0102F <sub>1</sub>	0102F <sub>1</sub>	-3.06(11) $\times 10^{-2}$
863	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	1(1, 0F <sub>1</sub> )	0102F <sub>1</sub>	0102F <sub>2</sub>	-2.973(56) $\times 10^{-2}$
875	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	0(0, 0A <sub>1</sub> )	0102F <sub>2</sub>	0102F <sub>2</sub>	-8.5(1.1) $\times 10^{-2}$
876	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	1(1, 0F <sub>1</sub> )	0102F <sub>2</sub>	0102F <sub>2</sub>	4.66(11) $\times 10^{-2}$
877	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	2(0, 0A <sub>1</sub> )	0102F <sub>2</sub>	0102F <sub>2</sub>	5.66(26) $\times 10^{-4}$

### Appendix B. Effective Hamiltonian parameters from the “ $\nu_1$ ” fit.

Table B.7 below gives the fitted values of the effective Hamiltonian parameters of  $\text{CF}_4$  for the “ $\nu_1$ ” fit (see Section 4.2). Notations correspond to Equation  
 655 (3). Standard deviation is given in parentheses in the unit of the last two digits.

Parameter	Polyad	Order	$\Omega(K, nC)$	$\{s\}$ $C_v$	$\{s'\}$ $C'_v$	Value / $\text{cm}^{-1}$
64	$\nu_1$	0	0(0, 0A <sub>1</sub> )	1000A <sub>1</sub>	1000A <sub>1</sub>	909.072146(52)
65	$\nu_1$	2	2(0, 0A <sub>1</sub> )	1000A <sub>1</sub>	1000A <sub>1</sub>	-3.47048(68) $\times 10^{-4}$
66	$\nu_1$	4	4(0, 0A <sub>1</sub> )	1000A <sub>1</sub>	1000A <sub>1</sub>	-2.65(30) $\times 10^{-10}$
67	$\nu_1$	4	4(4, 0A <sub>1</sub> )	1000A <sub>1</sub>	1000A <sub>1</sub>	-1.31(30) $\times 10^{-11}$
68	$\nu_1$	6	6(0, 0A <sub>1</sub> )	1000A <sub>1</sub>	1000A <sub>1</sub>	6.51(38) $\times 10^{-14}$
69	$\nu_1$	6	6(4, 0A <sub>1</sub> )	1000A <sub>1</sub>	1000A <sub>1</sub>	-5.47(28) $\times 10^{-15}$
70	$\nu_1$	6	6(6, 0A <sub>1</sub> )	1000A <sub>1</sub>	1000A <sub>1</sub>	1.186(74) $\times 10^{-15}$
299	$\nu_1 + \nu_4$	2	0(0, 0A <sub>1</sub> )	1001F <sub>2</sub>	1001F <sub>2</sub>	-6.7356(23) $\times 10^{-1}$
300	$\nu_1 + \nu_4$	3	1(1, 0F <sub>1</sub> )	1001F <sub>2</sub>	1001F <sub>2</sub>	6.547(31) $\times 10^{-3}$
301	$\nu_1 + \nu_4$	4	2(0, 0A <sub>1</sub> )	1001F <sub>2</sub>	1001F <sub>2</sub>	-7.0(1.2) $\times 10^{-6}$
302	$\nu_1 + \nu_4$	4	2(2, 0E )	1001F <sub>2</sub>	1001F <sub>2</sub>	1.31(17) $\times 10^{-5}$
303	$\nu_1 + \nu_4$	4	2(2, 0F <sub>2</sub> )	1001F <sub>2</sub>	1001F <sub>2</sub>	2.25(25) $\times 10^{-5}$
304	$\nu_1 + \nu_4$	5	3(1, 0F <sub>1</sub> )	1001F <sub>2</sub>	1001F <sub>2</sub>	1.49(24) $\times 10^{-7}$
305	$\nu_1 + \nu_4$	5	3(3, 0F <sub>1</sub> )	1001F <sub>2</sub>	1001F <sub>2</sub>	2.78(14) $\times 10^{-7}$
306	$\nu_1 + \nu_4$	6	4(0, 0A <sub>1</sub> )	1001F <sub>2</sub>	1001F <sub>2</sub>	1.86(18) $\times 10^{-8}$
307	$\nu_1 + \nu_4$	6	4(2, 0E )	1001F <sub>2</sub>	1001F <sub>2</sub>	4.96(19) $\times 10^{-8}$
308	$\nu_1 + \nu_4$	6	4(2, 0F <sub>2</sub> )	1001F <sub>2</sub>	1001F <sub>2</sub>	-1.55(29) $\times 10^{-8}$
309	$\nu_1 + \nu_4$	6	4(4, 0A <sub>1</sub> )	1001F <sub>2</sub>	1001F <sub>2</sub>	-1.83(17) $\times 10^{-9}$
310	$\nu_1 + \nu_4$	6	4(4, 0E )	1001F <sub>2</sub>	1001F <sub>2</sub>	-6.29(27) $\times 10^{-8}$
311	$\nu_1 + \nu_4$	6	4(4, 0F <sub>2</sub> )	1001F <sub>2</sub>	1001F <sub>2</sub>	-4.87(20) $\times 10^{-8}$
312	$\nu_1 + \nu_4$	7	5(1, 0F <sub>1</sub> )	1001F <sub>2</sub>	1001F <sub>2</sub>	3.55(43) $\times 10^{-11}$
313	$\nu_1 + \nu_4$	7	5(3, 0F <sub>1</sub> )	1001F <sub>2</sub>	1001F <sub>2</sub>	3.29(42) $\times 10^{-11}$
315	$\nu_1 + \nu_4$	7	5(5, 1F <sub>1</sub> )	1001F <sub>2</sub>	1001F <sub>2</sub>	-1.43(19) $\times 10^{-11}$
316	$\nu_1 + \nu_4$	8	6(0, 0A <sub>1</sub> )	1001F <sub>2</sub>	1001F <sub>2</sub>	-1.143(82) $\times 10^{-11}$
317	$\nu_1 + \nu_4$	8	6(2, 0E )	1001F <sub>2</sub>	1001F <sub>2</sub>	1.372(26) $\times 10^{-11}$
318	$\nu_1 + \nu_4$	8	6(2, 0F <sub>2</sub> )	1001F <sub>2</sub>	1001F <sub>2</sub>	-5.30(60) $\times 10^{-12}$
320	$\nu_1 + \nu_4$	8	6(4, 0E )	1001F <sub>2</sub>	1001F <sub>2</sub>	-2.337(46) $\times 10^{-11}$
321	$\nu_1 + \nu_4$	8	6(4, 0F <sub>2</sub> )	1001F <sub>2</sub>	1001F <sub>2</sub>	-1.320(61) $\times 10^{-11}$
322	$\nu_1 + \nu_4$	8	6(6, 0A <sub>1</sub> )	1001F <sub>2</sub>	1001F <sub>2</sub>	-6.82(85) $\times 10^{-13}$
323	$\nu_1 + \nu_4$	8	6(6, 0E )	1001F <sub>2</sub>	1001F <sub>2</sub>	4.47(18) $\times 10^{-12}$

Table B.7: Effective Hamiltonian parameters.

Table B.7 (continued).

Parameter	Polyad	Order	$\Omega(K, nC)$	$\{s\}$ $C_v$	$\{s'\}$ $C'_v$	Value / $\text{cm}^{-1}$
325	$\nu_1 + \nu_4$	8	6(6, 1F <sub>2</sub> )	1001F <sub>2</sub>	1001F <sub>2</sub>	5.99(24) $\times 10^{-12}$
326	$2\nu_1/\nu_3 + \nu_4$	2	0(0, 0A <sub>1</sub> )	2000A <sub>1</sub>	2000A <sub>1</sub>	-2.328912(48)
327	$2\nu_1/\nu_3 + \nu_4$	4	2(0, 0A <sub>1</sub> )	2000A <sub>1</sub>	2000A <sub>1</sub>	-3.252(20) $\times 10^{-6}$
329	$2\nu_1/\nu_3 + \nu_4$	6	4(4, 0A <sub>1</sub> )	2000A <sub>1</sub>	2000A <sub>1</sub>	-1.82(16) $\times 10^{-11}$
540	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	2	0(0, 0A <sub>1</sub> )	1010F <sub>2</sub>	1010F <sub>2</sub>	-7.268(32)
541	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	3	1(1, 0F <sub>1</sub> )	1010F <sub>2</sub>	1010F <sub>2</sub>	2.54(12) $\times 10^{-2}$
542	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	2(0, 0A <sub>1</sub> )	1010F <sub>2</sub>	1010F <sub>2</sub>	1.03(14) $\times 10^{-4}$
543	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	2(2, 0E)	1010F <sub>2</sub>	1010F <sub>2</sub>	-1.068(14) $\times 10^{-3}$
544	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	2(2, 0F <sub>2</sub> )	1010F <sub>2</sub>	1010F <sub>2</sub>	-1.79(20) $\times 10^{-4}$
545	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	3(1, 0F <sub>1</sub> )	1010F <sub>2</sub>	1010F <sub>2</sub>	-1.358(33) $\times 10^{-5}$
546	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	3(3, 0F <sub>1</sub> )	1010F <sub>2</sub>	1010F <sub>2</sub>	1.775(27) $\times 10^{-5}$
547	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	4(0, 0A <sub>1</sub> )	1010F <sub>2</sub>	1010F <sub>2</sub>	-3.591(66) $\times 10^{-7}$
550	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	4(4, 0A <sub>1</sub> )	1010F <sub>2</sub>	1010F <sub>2</sub>	-2.11(11) $\times 10^{-8}$
552	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	4(4, 0F <sub>2</sub> )	1010F <sub>2</sub>	1010F <sub>2</sub>	5.99(24) $\times 10^{-8}$
553	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	5(1, 0F <sub>1</sub> )	1010F <sub>2</sub>	1010F <sub>2</sub>	-9.85(45) $\times 10^{-10}$
554	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	5(3, 0F <sub>1</sub> )	1010F <sub>2</sub>	1010F <sub>2</sub>	9.26(43) $\times 10^{-10}$
556	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	5(5, 1F <sub>1</sub> )	1010F <sub>2</sub>	1010F <sub>2</sub>	-2.273(62) $\times 10^{-9}$
557	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	6(0, 0A <sub>1</sub> )	1010F <sub>2</sub>	1010F <sub>2</sub>	-4.50(15) $\times 10^{-11}$
567	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	2(2, 0F <sub>2</sub> )	1010F <sub>2</sub>	1002A <sub>1</sub>	-1.451(30) $\times 10^{-3}$
568	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	3(3, 0F <sub>2</sub> )	1010F <sub>2</sub>	1002A <sub>1</sub>	2.199(38) $\times 10^{-5}$
569	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	4(2, 0F <sub>2</sub> )	1010F <sub>2</sub>	1002A <sub>1</sub>	2.122(48) $\times 10^{-7}$
573	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	1(1, 0F <sub>1</sub> )	1010F <sub>2</sub>	1002E	-1.205(21) $\times 10^{-1}$
574	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	2(2, 0F <sub>2</sub> )	1010F <sub>2</sub>	1002E	-1.692(36) $\times 10^{-3}$
575	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	3(1, 0F <sub>1</sub> )	1010F <sub>2</sub>	1002E	-1.313(29) $\times 10^{-5}$
579	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	4(4, 0F <sub>1</sub> )	1010F <sub>2</sub>	1002E	-1.14(39) $\times 10^{-8}$
580	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	4(4, 0F <sub>2</sub> )	1010F <sub>2</sub>	1002E	-9.22(22) $\times 10^{-8}$
587	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	3	0(0, 0A <sub>1</sub> )	1010F <sub>2</sub>	1002F <sub>2</sub>	2.259(43)
588	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	1(1, 0F <sub>1</sub> )	1010F <sub>2</sub>	1002F <sub>2</sub>	-4.52(16) $\times 10^{-2}$
589	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	2(0, 0A <sub>1</sub> )	1010F <sub>2</sub>	1002F <sub>2</sub>	-1.401(23) $\times 10^{-3}$
590	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	2(2, 0E)	1010F <sub>2</sub>	1002F <sub>2</sub>	2.480(35) $\times 10^{-3}$



Table B.7 (continued).

Parameter	Polyad	Order	$\Omega(K, nC)$	$\{s\} C_v$	$\{s'\} C'_v$	Value / $\text{cm}^{-1}$
591	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	2(2, 0F <sub>2</sub> )	1010F <sub>2</sub>	1002F <sub>2</sub>	-7.94(33) $\times 10^{-4}$
592	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	3(1, 0F <sub>1</sub> )	1010F <sub>2</sub>	1002F <sub>2</sub>	1.717(53) $\times 10^{-5}$
593	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	3(3, 0F <sub>1</sub> )	1010F <sub>2</sub>	1002F <sub>2</sub>	-1.957(31) $\times 10^{-5}$
594	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	3(3, 0F <sub>2</sub> )	1010F <sub>2</sub>	1002F <sub>2</sub>	-3.544(68) $\times 10^{-5}$
595	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	4(0, 0A <sub>1</sub> )	1010F <sub>2</sub>	1002F <sub>2</sub>	4.305(67) $\times 10^{-7}$
596	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	4(2, 0E)	1010F <sub>2</sub>	1002F <sub>2</sub>	2.412(41) $\times 10^{-7}$
598	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	4(4, 0A <sub>1</sub> )	1010F <sub>2</sub>	1002F <sub>2</sub>	-2.49(23) $\times 10^{-8}$
599	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	4(4, 0E)	1010F <sub>2</sub>	1002F <sub>2</sub>	-7.08(30) $\times 10^{-8}$
600	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	4(4, 0F <sub>1</sub> )	1010F <sub>2</sub>	1002F <sub>2</sub>	-6.44(26) $\times 10^{-8}$
602	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	5(1, 0F <sub>1</sub> )	1010F <sub>2</sub>	1002F <sub>2</sub>	2.459(55) $\times 10^{-9}$
603	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	5(3, 0F <sub>1</sub> )	1010F <sub>2</sub>	1002F <sub>2</sub>	-3.061(67) $\times 10^{-9}$
604	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	5(3, 0F <sub>2</sub> )	1010F <sub>2</sub>	1002F <sub>2</sub>	9.47(88) $\times 10^{-10}$
606	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	5(5, 0F <sub>1</sub> )	1010F <sub>2</sub>	1002F <sub>2</sub>	1.627(43) $\times 10^{-9}$
607	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	5(5, 1F <sub>1</sub> )	1010F <sub>2</sub>	1002F <sub>2</sub>	9.73(36) $\times 10^{-10}$
608	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	5(5, 0F <sub>2</sub> )	1010F <sub>2</sub>	1002F <sub>2</sub>	1.110(30) $\times 10^{-9}$
609	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	0(0, 0A <sub>1</sub> )	1002A <sub>1</sub>	1002A <sub>1</sub>	7.4(1.9) $\times 10^{-1}$
610	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	2(0, 0A <sub>1</sub> )	1002A <sub>1</sub>	1002A <sub>1</sub>	6.50(24) $\times 10^{-3}$
611	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	4(0, 0A <sub>1</sub> )	1002A <sub>1</sub>	1002A <sub>1</sub>	-2.915(79) $\times 10^{-6}$
613	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	2(2, 0E)	1002A <sub>1</sub>	1002E	-2.50(20) $\times 10^{-4}$
614	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	4(2, 0E)	1002A <sub>1</sub>	1002E	-1.139(52) $\times 10^{-7}$
615	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	4(4, 0E)	1002A <sub>1</sub>	1002E	1.228(27) $\times 10^{-7}$
616	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	2(2, 0F <sub>2</sub> )	1002A <sub>1</sub>	1002F <sub>2</sub>	-1.367(42) $\times 10^{-3}$
617	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	3(3, 0F <sub>2</sub> )	1002A <sub>1</sub>	1002F <sub>2</sub>	-1.495(44) $\times 10^{-5}$
618	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	4(2, 0F <sub>2</sub> )	1002A <sub>1</sub>	1002F <sub>2</sub>	-3.817(99) $\times 10^{-7}$
619	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	4(4, 0F <sub>2</sub> )	1002A <sub>1</sub>	1002F <sub>2</sub>	2.928(29) $\times 10^{-7}$
620	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	0(0, 0A <sub>1</sub> )	1002E	1002E	-3.15(21) $\times 10^{-2}$
621	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	2(0, 0A <sub>1</sub> )	1002E	1002E	2.36(17) $\times 10^{-4}$
622	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	2(2, 0E)	1002E	1002E	1.472(22) $\times 10^{-3}$
623	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	3(3, 0A <sub>2</sub> )	1002E	1002E	2.76(23) $\times 10^{-6}$
625	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	4(2, 0E)	1002E	1002E	2.89(38) $\times 10^{-8}$

Table B.7 (continued).

Parameter	Polyad	Order	$\Omega(K, nC)$	$\{s\}$ $C_v$	$\{s'\}$ $C'_v$	Value / $\text{cm}^{-1}$
626	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	4(4, 0A <sub>1</sub> )	1002E	1002E	4.08(15) $\times 10^{-8}$
627	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	4(4, 0E)	1002E	1002E	-4.135(29) $\times 10^{-7}$
628	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	1(1, 0F <sub>1</sub> )	1002E	1002F <sub>2</sub>	-5.35(11) $\times 10^{-2}$
629	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	2(2, 0F <sub>2</sub> )	1002E	1002F <sub>2</sub>	-3.61(18) $\times 10^{-4}$
630	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	3(1, 0F <sub>1</sub> )	1002E	1002F <sub>2</sub>	-1.741(26) $\times 10^{-5}$
631	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	3(3, 0F <sub>1</sub> )	1002E	1002F <sub>2</sub>	2.512(22) $\times 10^{-5}$
632	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	3(3, 0F <sub>2</sub> )	1002E	1002F <sub>2</sub>	-1.248(27) $\times 10^{-5}$
633	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	4(2, 0F <sub>2</sub> )	1002E	1002F <sub>2</sub>	1.136(33) $\times 10^{-7}$
634	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	4(4, 0F <sub>1</sub> )	1002E	1002F <sub>2</sub>	3.048(27) $\times 10^{-7}$
635	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	4(4, 0F <sub>2</sub> )	1002E	1002F <sub>2</sub>	9.13(35) $\times 10^{-8}$
636	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	4	0(0, 0A <sub>1</sub> )	1002F <sub>2</sub>	1002F <sub>2</sub>	1.243(32)
637	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	5	1(1, 0F <sub>1</sub> )	1002F <sub>2</sub>	1002F <sub>2</sub>	-3.38(12) $\times 10^{-2}$
638	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	2(0, 0A <sub>1</sub> )	1002F <sub>2</sub>	1002F <sub>2</sub>	-8.03(15) $\times 10^{-4}$
639	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	2(2, 0E)	1002F <sub>2</sub>	1002F <sub>2</sub>	2.118(18) $\times 10^{-3}$
640	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	6	2(2, 0F <sub>2</sub> )	1002F <sub>2</sub>	1002F <sub>2</sub>	-3.82(26) $\times 10^{-4}$
641	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	7	3(1, 0F <sub>1</sub> )	1002F <sub>2</sub>	1002F <sub>2</sub>	5.23(34) $\times 10^{-6}$
645	$\nu_2 + \nu_3/\nu_2 + 2\nu_4$	8	4(2, 0F <sub>2</sub> )	1002F <sub>2</sub>	1002F <sub>2</sub>	-5.002(49) $\times 10^{-7}$